

Cosmic
Strings

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Challenge

Cosmic
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Black Holes

Globular

Conclusions

Supermassive Black Holes from Superconducting Cosmic Strings

Robert Brandenberger
McGill University

Workshop in Honour of Brandon Carter, IAP, July 4 - 6 2022

Work in collaboration with B. Cyr and H. Jiao,
[arXiv:2202.01799](https://arxiv.org/abs/2202.01799)

What I am not talking about

S. Brahma, RB, S. Laliberte, arXiv:2206.12468

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Emergent metric space-time from BFSS matrix model

- Starting point: **BFSS** matrix model in a high temperature state.
- Theory of 10 Hermitean $N \times N$ matrices A_μ , $N \rightarrow \infty$ limit.
- Emergent continuous and infinite time** from A_0 matrix.
- Emergent continuous and infinite space** from A_i matrices
- Phase transition: $S_0(9) \rightarrow SO(3) \times SO(6)$, only three dimensions become macroscopic.
- Emergent spatial metric** $g_{ij}(t) = A(t)\delta_{ij}$: spatially flat.
- Thermal fluctuations \rightarrow scale-invariant spectra of density fluctuations and gravitational waves. (S. Brahma, RB and S. Laliberte, arXiv:2107.11512).

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Supermassive black holes from superconducting cosmic strings

B. Cyr, H. Jiao and RB, arXiv:2202.01799, MNRAS, in press

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- Black Holes: one focal point of Brandon's research.
- Superconducting cosmic strings: another focal point of Brandon's research.
- This talk: Combination of these two interests of Brandon.
- Loops of superconducting cosmic strings can seed direct collapse black hole formation at high redshifts.
- → explanation for the origin and abundance of observed high redshift super-massive black holes.

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Outline

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- 1 Observational Challenge
- 2 Cosmic String Review
- 3 Super-Massive High Redshift Black Holes from Superconducting Cosmic Strings
- 4 Globular Clusters from Cosmic Strings
- 5 Conclusions

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High Redshift Super-Massive Black Holes: Challenge for Standard Λ CDM Paradigm

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- Black holes with masses $M > 10^9 M_{\odot}$ observed at redshifts $z > 6$.
- Accretion bounded by Eddington rate.
- → high mass nonlinear seeds required at early times.
- Standard Λ CDM model: probability of such nonlinear seeds exponentially suppressed.

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Required Seed Mass (Eddington Accretion)

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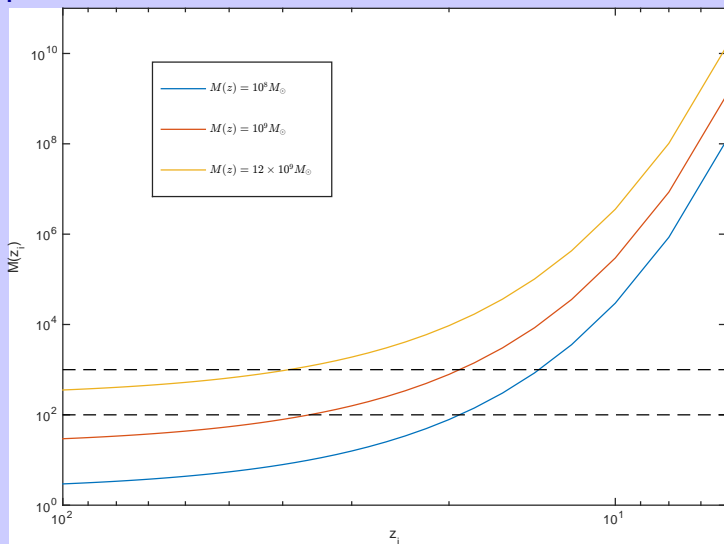
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Abundance of nonlinear overdensities in standard Λ CDM model

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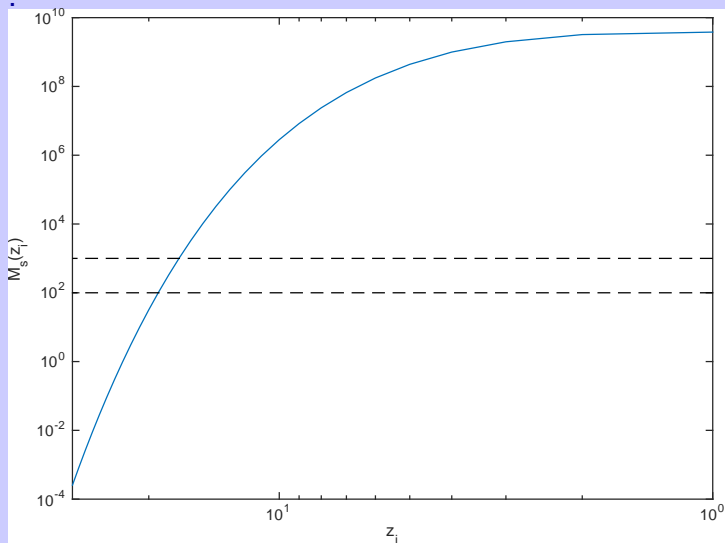
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Cosmic Strings to the Rescue

T. Kibble, J. Phys. A **9**, 1387 (1976); Y. B. Zeldovich, Mon. Not. Roy. Astron. Soc. **192**, 663 (1980); A. Vilenkin, Phys. Rev. Lett. **46**, 1169 (1981).

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- Assume: theory which describes our matter has **cosmic string solutions**.
- → scaling distribution of strings at all times.
- **Cosmic string loops** → **nonlinear perturbations at high redshifts**.
- → more massive seeds which have more time to grow.
- → solution of the supermassive black hole mystery.

Abundance of nonlinear overdensities due to cosmic strings

S. Bramberger, R.B., P. Jreidnin and J. Quintin, arXiv:1503.02317

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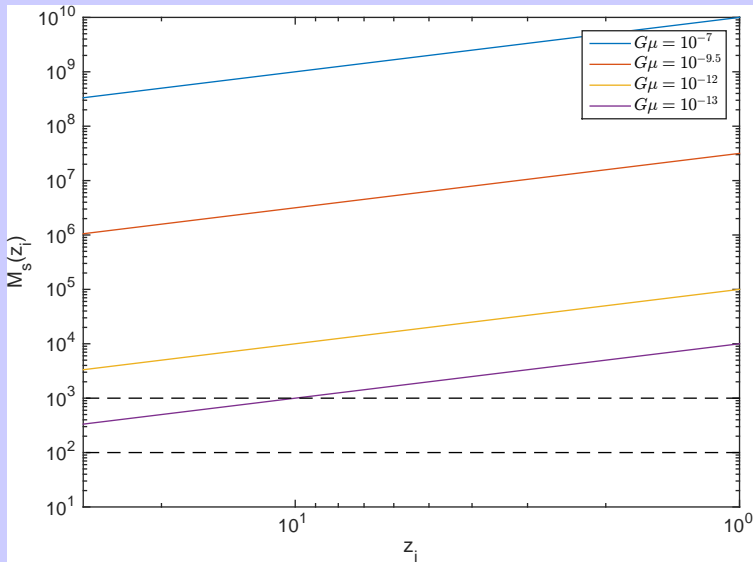
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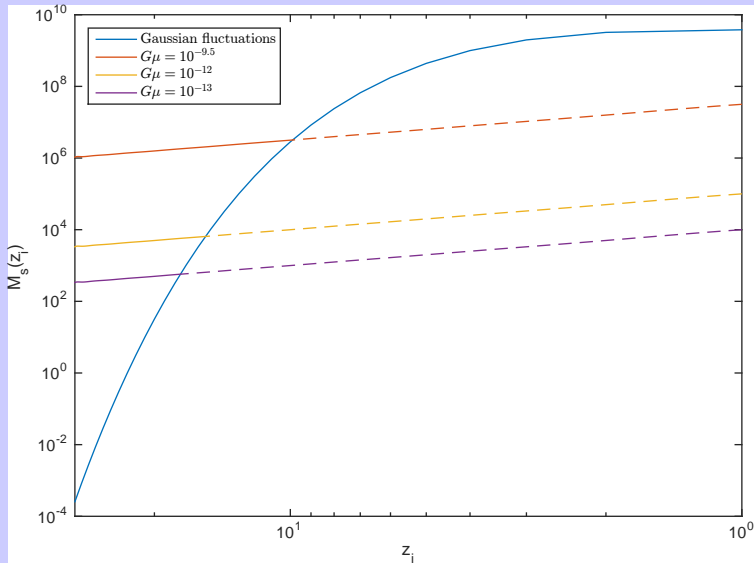
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New Challenge

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Conclusions

- Nonlinear seeds of sufficient mass is a necessary but not a sufficient criterion for black hole formation.
- The mass needs to collapse to within its Schwarzschild radius.
- In general a collapsing cloud will fragment → no black hole formation.
- Presence of Lyman-Werner radiation can prevent the fragmentation.
- Superconducting cosmic strings produce Lyman-Werner radiation.
- Superconducting cosmic string loops → direct collapse black hole formation.

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- **Superconducting cosmic strings** produce Lyman-Werner radiation.
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Conclusions

- **Cosmic string = linear topological defect** in a quantum field theory.
- 1st analog: line defect in a crystal
- 2nd analog: vortex line in superfluid or superconductor
- **Cosmic string = line of trapped energy density** in a quantum field theory.
- Trapped energy density \rightarrow gravitational effects on space-time \rightarrow important in cosmology.

Relevance to Particle Physics I

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Conclusions

- Cosmic string solutions **exist** in many particle physics models **beyond the “Standard Model”**.
- In models which admit cosmic strings, cosmic strings **inevitably form** in the early universe and **persist to the present time**.
- Seeing a cosmic string in the sky would provide a guide to particle physics beyond the Standard Model!
- Cosmic strings approach a **scaling solution**: the statistical properties of the string network is independent of time if all length are scaled to the Hubble radius.
- Scaling solution is independent of the string tension.
- String network consists of **long strings** and **string loops**.

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Relevance to Particle Physics II

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Conclusions

- Cosmic strings are characterized by their **tension** μ which is associated with the energy scale η at which the strings form ($\mu \sim \eta^2$).
- Searching for the signatures of cosmic strings is a **tool to probe physics beyond the Standard Model** at energy ranges complementary to those probed by the LHC.
- Cosmic strings are constrained from cosmology: $G\mu \leq 1.3 \times 10^{-7}$ otherwise a conflict with the observed acoustic oscillations in the CMB angular power spectrum (Dvorkin, Hu and Wyman, 2011).
 - String loops oscillate and emit gravitational waves.
 - Constraints as low as $G\mu < 10^{-10}$ follow from pulsar timing constraints on the amplitude of the spectrum of stochastic GW background.
 - Existing **upper bound** on the string tension rules out large classes of “Grand Unified” models.

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Relevance to Cosmology

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Conclusions

Strings can produce many **good things** for cosmology:

- String-induced mechanism of baryogenesis (R.B., A-C. Davis and M. Hindmarsh, 1991).
- Explanation for the origin of primordial magnetic fields which are coherent on galactic scales (X.Zhang and R.B. (1999)).
- **Seeds for high redshift supermassive black holes** (S. Bramberger, R.B., P. Jreidini and J. Quintin, 2015; R.B., B. Cyr and H. Jiao, 2021).
- **Origin of globular clusters** (A. Barton, R.B. and L. Lin, 2015; R.B., L. Lin and S. Yamanouchi, 2015).
- Origin of fast radio bursts (R.B., B. Cyr and A. Iyer, 2017).
- Global 21-cm absorption signal (EDGES) (R. Thériault, J. Mirocha and R.B. 2021)

Cosmic string scaling solution

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Conclusions

- At any time t there will be a network of “infinite” strings with curvature radius $\sim t$, ~ 1 per Hubble volume.
- At any time t there will be a distribution of string loops with radii R

$$n(R, t) = NR^{-5/2}t_{eq}^{1/2}t^{-2} \quad \gamma G\mu t < R < \alpha t$$

- $R_c = \gamma G\mu t$: gravitational radiation cutoff. Loops with $R < R_c$ negligible.

Superconducting Strings

E. Witten, Nucl. Phys. B249, 557 (1985)

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Conclusions

- In many models, cosmic strings can carry currents → **superconducting cosmic strings**.
- Bosonic and Fermionic superconductivity.
- In Grand Unified Models cosmic string superconductivity is generic.
- In superconducting string models there is an additional parameter characterizing the string network, the **current I** .
- “Maximal” current I_c : for $I > I_c$ electromagnetic radiation dominates → cutoff R_c on loop distribution changes..

Note: **Brandon Carter** pioneered a large body of work on superconducting cosmic strings, e.g. on **vorton** formation (B. Carter, Int. J. Theor. Phys. 36, 2451 (1997)).

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- **Primordial black holes:** Hubble scale nonlinearities form a black hole because the Schwarzschild radius equals the radius of the overdensity.
- Λ CDM model of cosmology \rightarrow nonlinearities form at late times and on scales much smaller than the Hubble radius. \rightarrow Schwarzschild radius is parametrically smaller than the radius of the overdensity..
- **Insufficient to have nonlinear fluctuations: Need to demonstrate that the mass collapses to inside the Schwarzschild radius.**
- In general, a collapsing gas cloud will fragment, form stars and never lead to a super-massive black hole (only stellar mass black holes).

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Direct Collapse Black Hole Criteria

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To allow a gas cloud to collapse into a super-massive black hole the following criteria must be satisfied:

- **Sufficient mass condition:** $M_b > 10^5 M_\odot$ to form a super-massive black hole.
- **Atomic cooling threshold condition:** Collapse without fragmentation $\rightarrow T_{vir} > 10^4 K$.
- **No heavy metal condition:** presence of heavy metals would allow cooling \rightarrow fragmentation.
- **No molecular hydrogen:** would lead to cooling and fragmentation \rightarrow requires presence of a **Lyman-Werner background** of $J > J_c \sim 10^{-44} \text{GeV}^3$.

Realizing the Direct Collapse Black Hole Criteria I

B. Cyr, H. Jiao and RB, arXiv:2202.01799, MNRAS, in press

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Sufficient mass condition at redshift $z < z_{rec}$:

$$M_b(z) = \frac{\Omega_b(z)}{\Omega_M(z)} \beta \mu R \frac{1 + z_{eq}}{1 + z} > 10^5 M_\odot$$

$$\rightarrow R_c < R < \alpha t_{eq}$$

There is a range of loop radii for which the condition is satisfied.

Atomic cooling condition:

Spherical collapse \rightarrow kinetic energy at collapse \rightarrow converted to virial temperature.

Result: atomic cooling condition satisfied whenever the mass condition is met.

Realizing the Direct Collapse Black Hole Criteria I

B. Cyr, H. Jiao and RB, arXiv:2202.01799, MNRAS, in press

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Realizing the Direct Collapse Black Hole Criteria II

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Lyman-Werner condition

$$\frac{dP}{d\omega} = \kappa I^2 R^{1/3} \omega^{-2/3}$$

Assumption: radiation remains confined in overdense region \rightarrow can compute the density of photons with $10\text{eV} < E < 13\text{eV}$

\rightarrow there is a range of currents $I < I_c$ for which the condition is satisfied.

Parameter Space Region

B. Cyr, H. Jiao and RB, arXiv:2202.01799, MNRAS, in press

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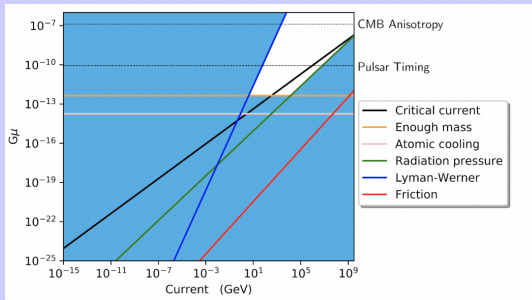
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Lessons

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Conclusions

There is a range of the cosmic string parameter space for which the direct collapse black hole criteria can be satisfied.

- For $G\mu \sim 10^{-10}$ the mean separation of loops forming SMBH will be $d_g \sim 10^{2/3} \text{Mpc}$
- \rightarrow reasonable number density of SMBH (M. Volonteri).

Intermediate Mass Black Holes from Cosmic String Loops?

RB, B. Cyr, and H. Jiao, arXiv:2103.14057

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Conclusions

- Consider $G\mu$ chosen to get the correct abundance of SMBH.
- $dn/dM \sim M^{-2}$ for $M_c < M < M_{SMBH} \rightarrow$ abundance of seeds which *might* lead to BH formation.
- $M_c \sim M_\odot$
- \rightarrow seeds in the *mass gap region* present in great abundance.

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Idea

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- **Globular clusters: oldest and most dense** star clusters, distributed in the **halo**.
- **Cosmic string loops: oldest and most dense** nonlinear regions, distributed in the **halo** of Λ CDM fluctuations.
- **Question:** Could string loops be the seeds of globular clusters?

Idea

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- **Question**: Could string loops be the seeds of globular clusters?

Abundance of globular cluster seeds due to cosmic strings

A. Barton, RB and L. Lin, arXiv:1502.07301

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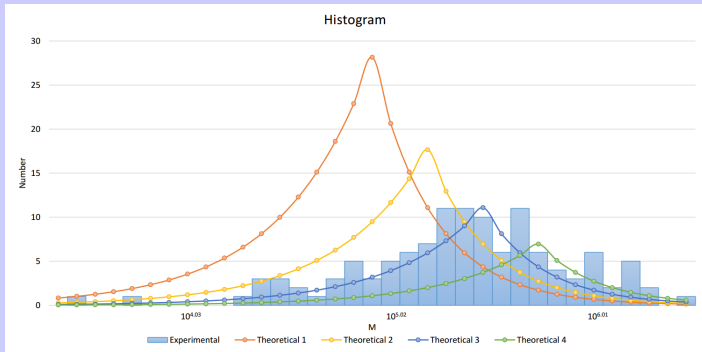
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Effect of Cosmic String Velocities

L. Lin, S. Yamanouchi and RB , arXiv:1508.02784

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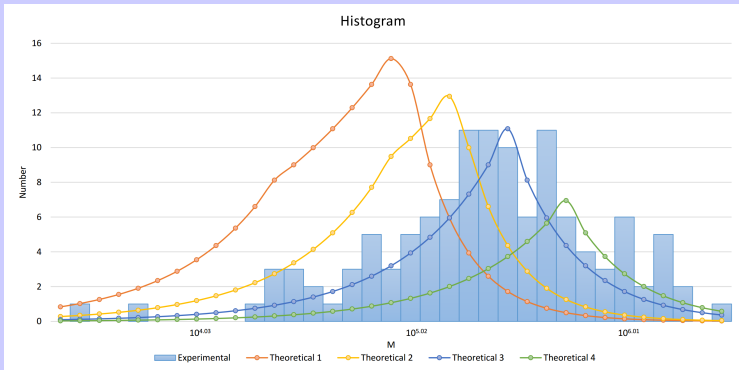
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- For $G\mu \sim 10^{-10}$ loops of superconducting cosmic strings can seed the observed abundance of high redshift super-massive black holes.
- Specifically: direct collapse black hole criteria can be satisfied in a range of cosmic string parameter space.
- String loops \rightarrow other interesting consequences for cosmology.