



# IUGONET activity for upper atmosphere study

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and IUGONET project team



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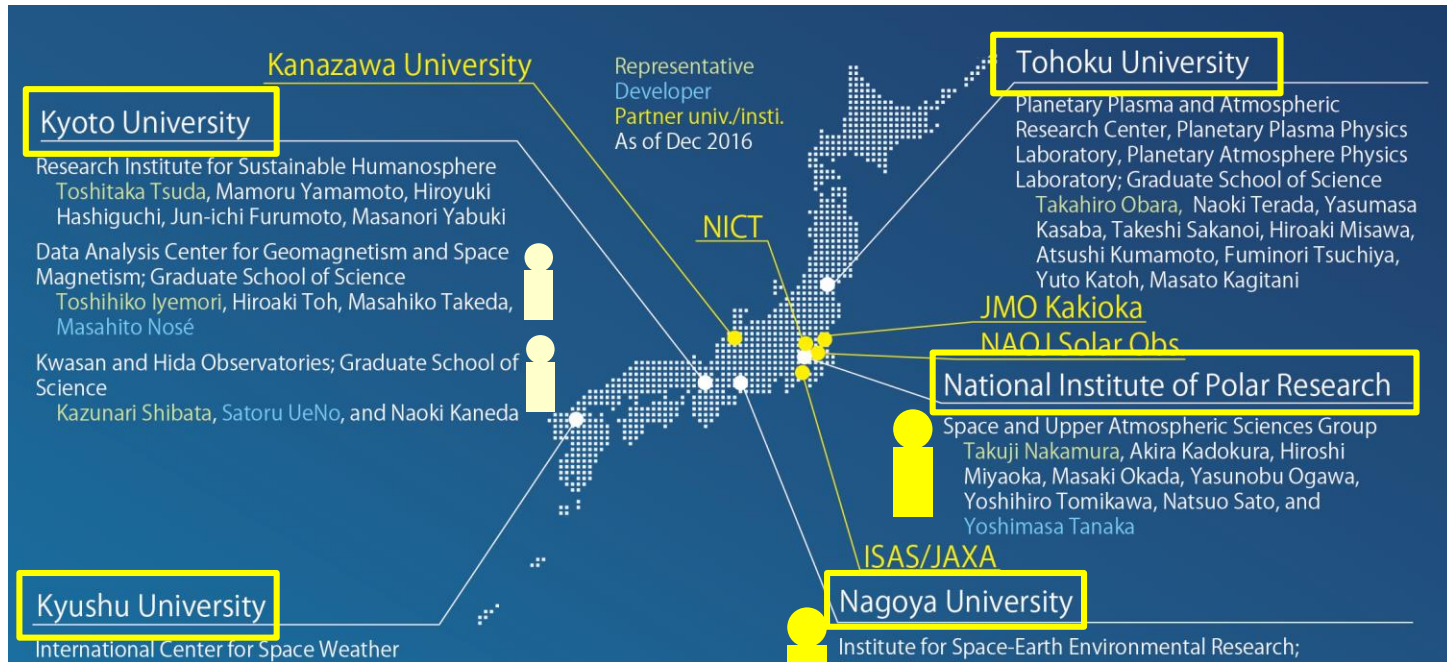
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<sup>6</sup> Kwasan and Hida Observatories, Graduate School of Science, Kyoto University.



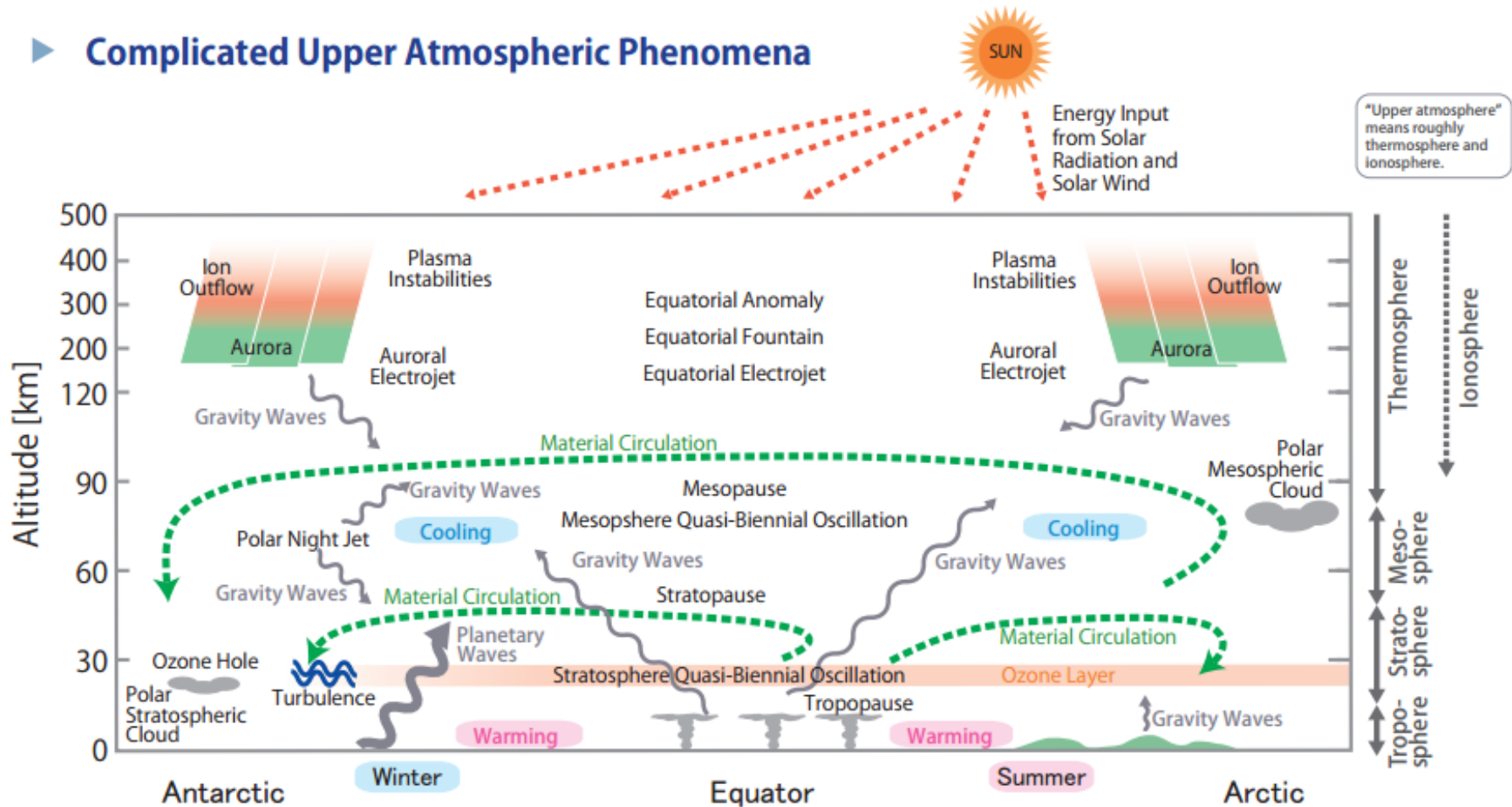
IUGONET (Inter-university Upper atmosphere Global Observation NETWORK)



IUGONET project started in 2009

- to **develop the tools for sharing the upper atmospheric data**, which have been archived separately by many Japanese universities and institutes since the International Geophysical Year (1957-1958).
- to comprehensively understand the mechanisms of long-term variations in the upper atmosphere and **facilitate interdisciplinary studies** regarding the Solar-Terrestrial Physics.

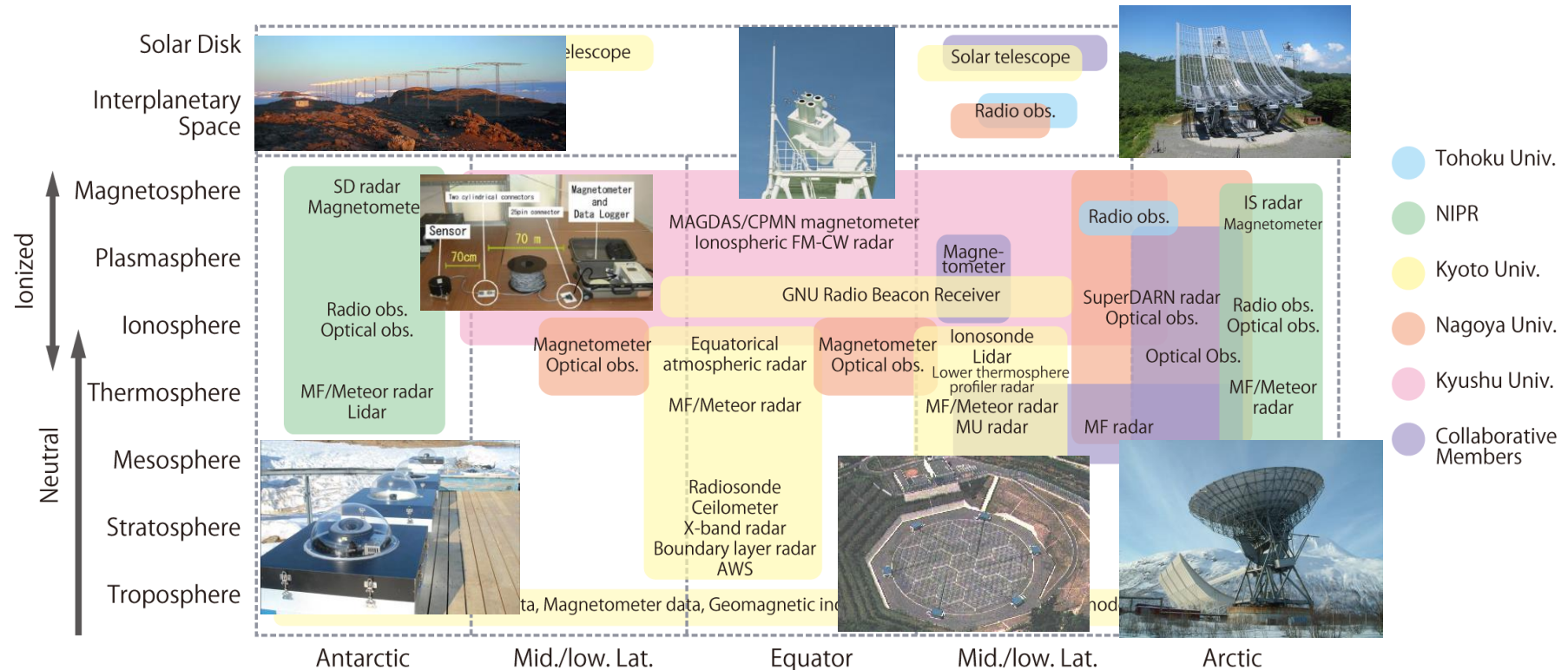
## ► Complicated Upper Atmospheric Phenomena



- Consists of multiple layers, such as the stratosphere, mesosphere, and thermosphere.
- Various phenomena are generated by energy inputs both from upper region (e.g., solar radiation, solar wind, energetic particle precipitation) and lower atmosphere (e.g., atmospheric waves).
- Both vertical coupling and horizontal circulation play an important role in the formation of the Earth's atmosphere.

# What data do we handle?

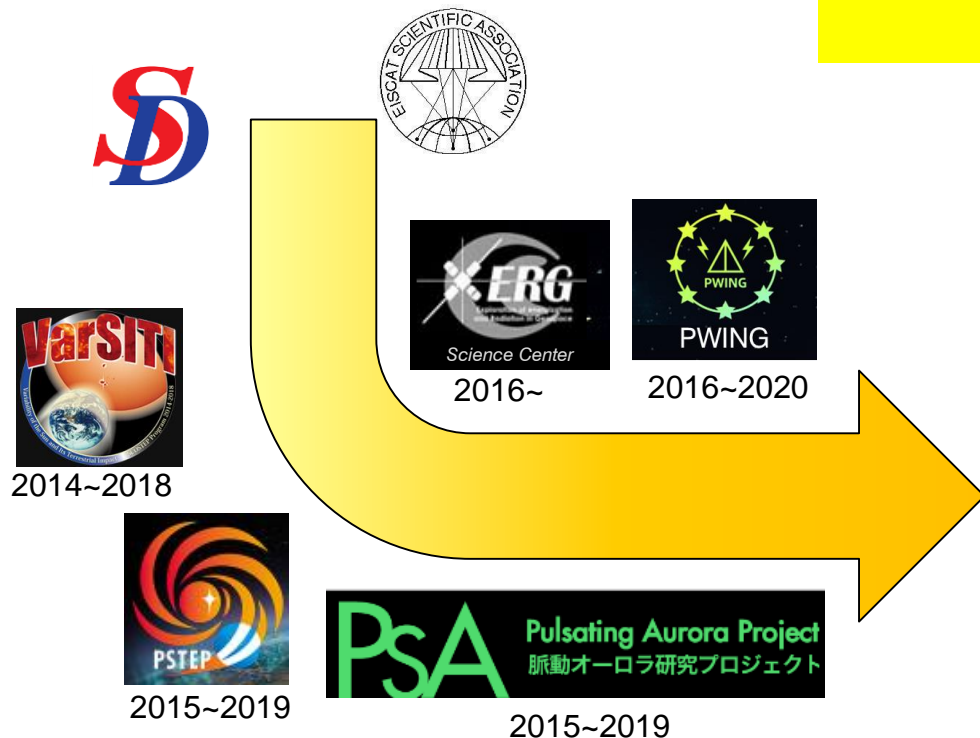
## ► IUGONET Global Network of Ground-Based Observations



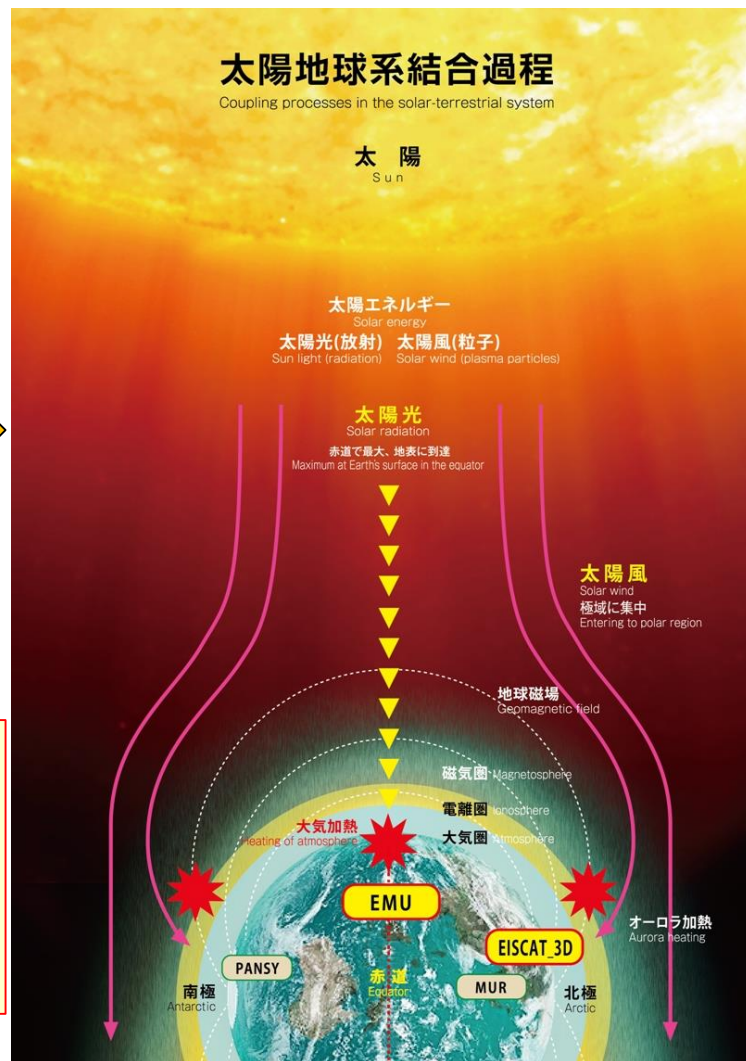
- IUGONET has handled data obtained by various kinds of instruments, such as telescopes, imagers, radars, and magnetometers, distributed globally all over the world.
- To understand the coupling processes in the solar-terrestrial system, it is important to comprehensively analyze these various kinds of data.



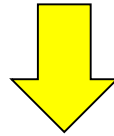
## Master Plan “Study of coupling processes in the solar-terrestrial system”



- We have collaborated with various projects associated with solar-terrestrial physics.
- We have supported publishing their data and have also provided useful tools for producing scientific results effectively.



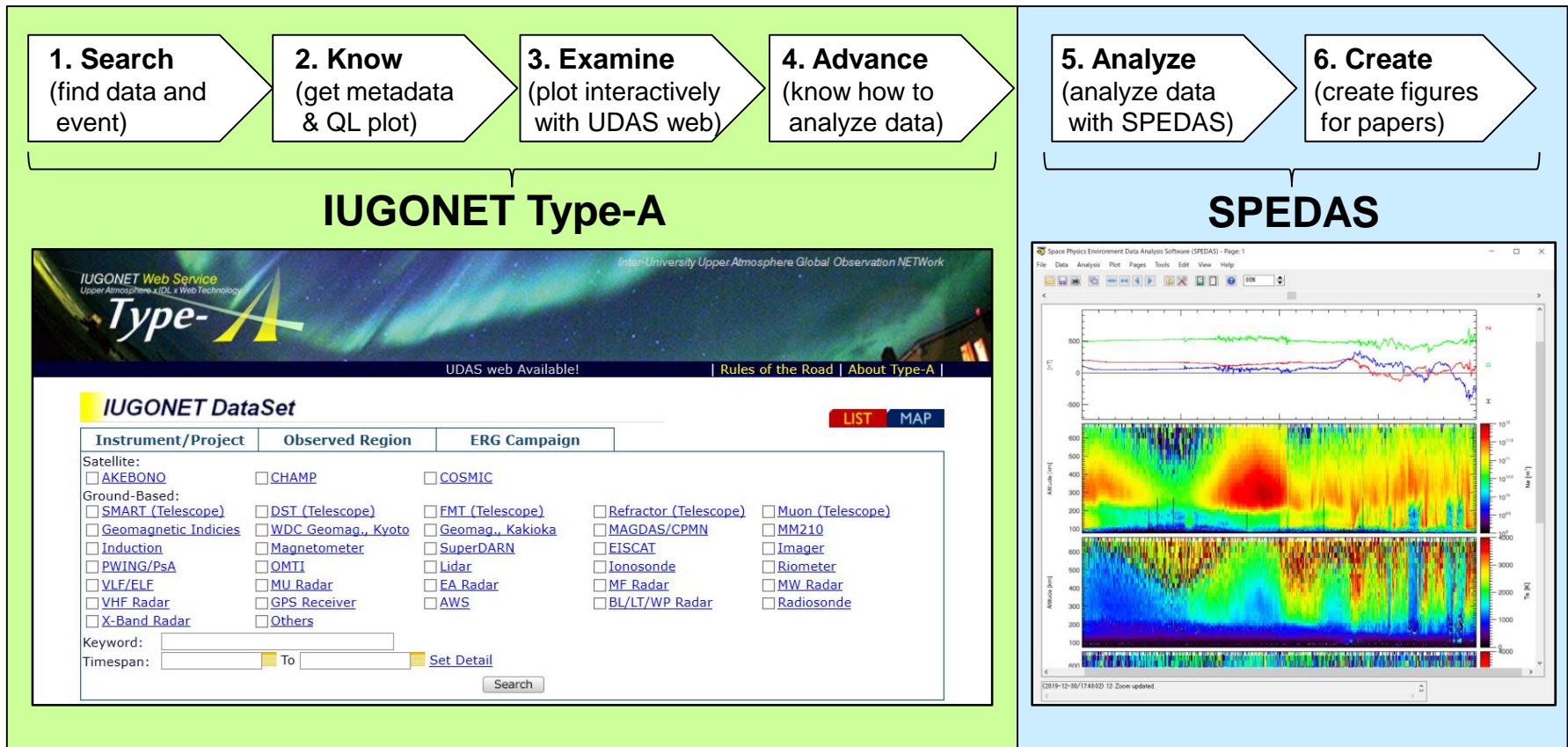
- How can we produce scientific results effectively?
- How can we make new discoveries from a lot of data?



## Our answers

- Connect seamlessly our research procedures (i.e., Search → Know (get metadata & QL plot) → Download → Visualize → Analyze data) using advanced tools.
- Remove barriers among various missions and communities.
- Contribute to the education of young researchers and promote international collaboration.

Our strategy is to help researchers perform comprehensive data analysis by connecting seamlessly research procedures using the advanced tools, i.e., **metadata database (IUGONET Type-A)** and **analysis software (SPEDAS)** for upper atmosphere data.





# IUGONET Type-A (1)

<http://search.iugonet.org/>

IUGONET Web Service Type-A

UDAS web Available! | Rules of the Road | About Type-A

### IUGONET DataSet

**Instrument/Project** **Observed Region** **ERG Campaign**

Satellite: ☐ AKEBONO ☐ CHAMP ☐ COSMIC

Ground-Based: ☐ SMART (Telescope) ☐ DST (Telescope) ☐ FMT (Telescope) ☐ Refractor (Telescope) ☐ Muon (Telescope)

☐ Geomagnetic Indices ☐ WDC Geomag., Kyoto ☐ Geomag., Kakioka ☐ MAGOAS/CPMN ☐ MM210

☐ Induction ☐ ONTI ☐ SuperDARN ☐ EISCAT ☐ Imager

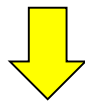
☐ PUVIS/PsA ☐ MU Radar ☐ Ionosonde ☐ Biometer

☐ VLF/ELF ☐ GPS Receiver ☐ AWS ☐ BL/LT/VP Radar ☐ MW Radar

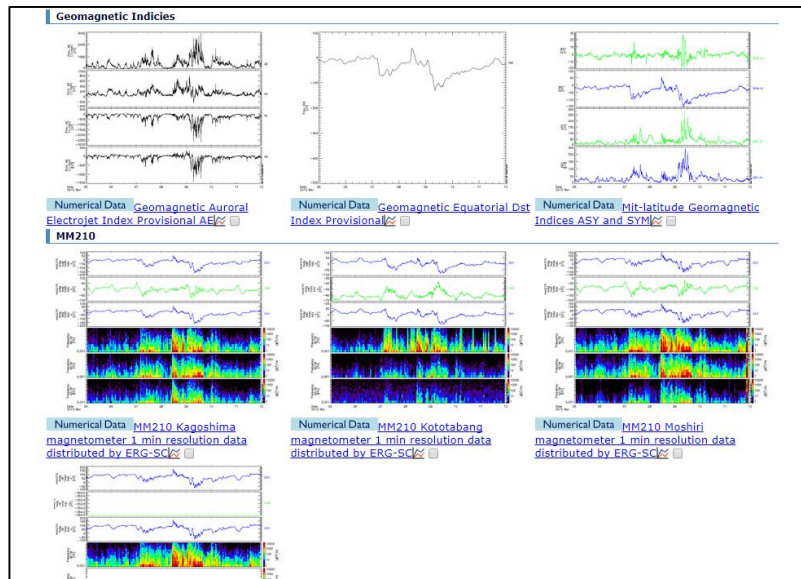
☐ VHF Radar ☐ Others

Keywords:

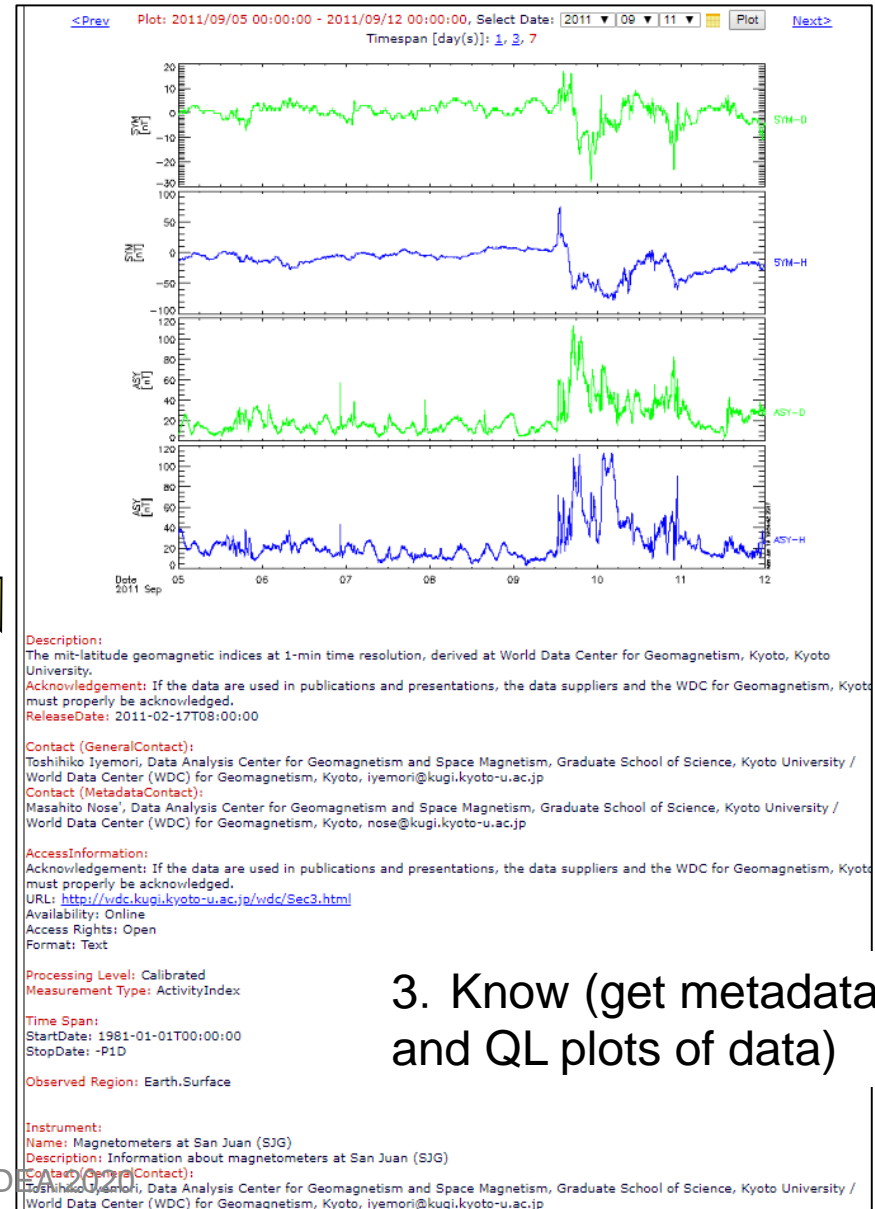
Timespan:  To



1. Search data



2. Find interesting events



3. Know (get metadata and QL plots of data)



The IUGONET metadata is based on the **Space Physics Archive Search and Extract (SPASE)** data model.

**Description:**

The mid-latitude geomagnetic indices at 1-min time resolution, derived at World Data Center for Geomagnetism, Kyoto, Kyoto University.

**Acknowledgement:** If the data are used in publications and presentations, the data suppliers and the WDC for Geomagnetism, Kyoto must properly be acknowledged.

**ReleaseDate:** 2011-02-17T08:00:00

**Contact (GeneralContact):**

Toshihiko Iyemori, Data Analysis Center for Geomagnetism and Space Magnetism, Graduate School of Science, Kyoto University / World Data Center (WDC) for Geomagnetism, Kyoto, iyemori@kugi.kyoto-u.ac.jp

**Contact (MetadataContact):**

Masahito Nose, Data Analysis Center for Geomagnetism and Space Magnetism, Graduate School of Science, Kyoto University / World Data Center (WDC) for Geomagnetism, Kyoto, nose@kugi.kyoto-u.ac.jp

**AccessInformation:**

Acknowledgement: If the data are used in publications and presentations, the data suppliers and the WDC for Geomagnetism, Kyoto must properly be acknowledged.

URL: <http://wdc.kugi.kyoto-u.ac.jp/wdc/Sec3.html>

Availability: Online

Access Rights: Open

Format: Text

**Processing Level:** Calibrated

**Measurement Type:** ActivityIndex

**Time Span:**

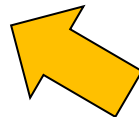
StartDate: 1981-01-01T00:00:00

StopDate: -P1D

**Observed Region:** Earth.Surface

3. Know (get metadata and  
QL plots of data)

## SPASE base model



<NumericalData>  
<Person>  
<Observatory>  
<Instrument>  
<Repository>

## UDAS web

### Step.1: Set Time Range

From: 2012 03 05 00 00 00  
To: 2012 03 12 00 00 00

### Step.2: Choose Variables to Plot

**Numerical Data** [MAGDAS observation network 1sec resolution](#)  
☒ magdas\_mag\_ama\_1sec\_hdz\_x (North-South magnetic field)  
☐ magdas\_mag\_ama\_1sec\_hdz\_y (East-West magnetic field)  
☐ magdas\_mag\_ama\_1sec\_hdz\_z (Vertical magnetic field)  
☐ magdas\_mag\_ama\_1sec\_hdz\_x\_dpwrspc (Dynamic power spectrum of North-South magnetic field)  
☐ magdas\_mag\_ama\_1sec\_hdz\_y\_dpwrspc (Dynamic power spectrum of East-West magnetic field)  
☐ magdas\_mag\_ama\_1sec\_hdz\_z\_dpwrspc (Dynamic power spectrum of Vertical magnetic field)

### **Numerical Data** [Geomagnetic Auroral Electrojet Index Provisional AE](#)

☒ wdc\_mag\_ae\_prov\_1min\_0 (Geomagnetic Auroral Electrojet (AE) index (AE = |AU - AL|))  
☐ wdc\_mag\_ae\_prov\_1min\_1 (Amplitude of Upper envelope (AU))  
☐ wdc\_mag\_ae\_prov\_1min\_2 (Amplitude of Lower envelope (AL))  
☐ wdc\_mag\_ae\_prov\_1min\_3 (Average value of the AU and AL induces (AO = (AU + AL)/2))  
☐ wdc\_mag\_ae\_prov\_1min\_4 (AX index)

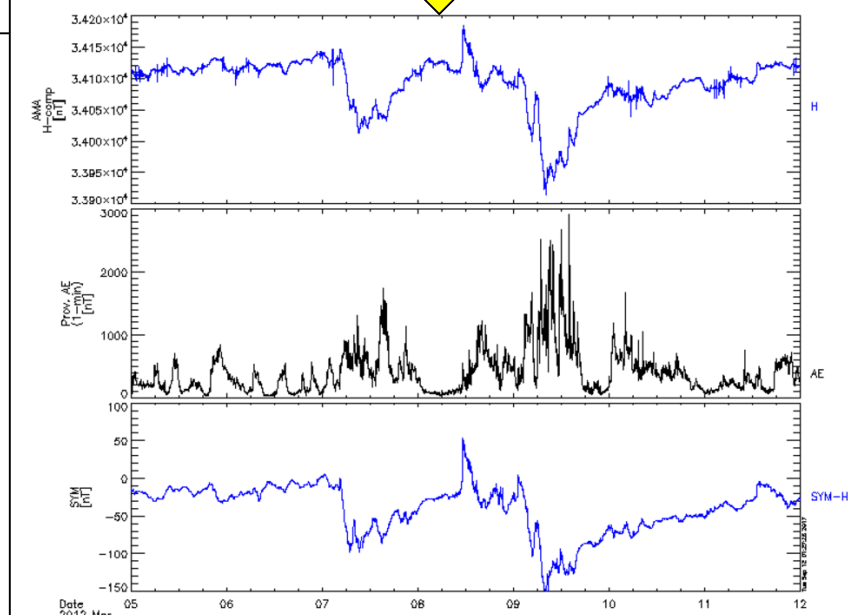
### **Numerical Data** [Mid-latitude Geomagnetic Indices ASY and SYM](#)

☒ wdc\_mag\_sym\_0 (Symetory (SYM) index of the east-west component)  
☒ wdc\_mag\_sym\_1 (Symetory (SYM) index of the north-south component)  
☐ wdc\_mag\_asy\_0 (Asymetory (ASY) index of the east-west component)  
☐ wdc\_mag\_asy\_1 (Asymetory (ASY) index of the north-south component)

\* At Least, one variable should be chosen.

Attentions: To create plot image requires some observation data, so read each acknowledgement(s) again.  
 (Expect if you want to see the plot only in this system)

4. Examine (plot data interactively with UDAS-web)



### How to Plot (SPEDAS-CUI #Basic):

```
IDL> thm_init
THEMIS> timespan, ['2012-03-05 00:00:00', '2012-03-12 00:00:00']
THEMIS> iug_load_gmag_wdc, site='sym asy'
THEMIS> tplot, ['wdc_mag_sym', 'wdc_mag_asy']
```

### How to Plot (SPEDAS-CUI #Advanced [\*Quick-Look was created with the command]):

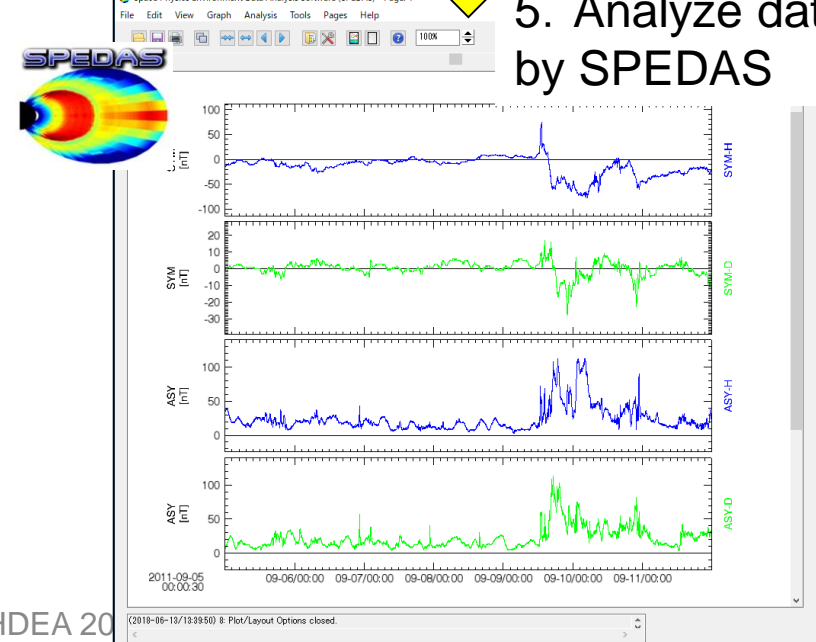
```
IDL> thm_init
THEMIS> timespan, ['2012-03-05 00:00:00', '2012-03-12 00:00:00']
THEMIS> iug_load_gmag_wdc, site='sym asy'
THEMIS> split_vec, 'wdc_mag_sym'
THEMIS> split_vec, 'wdc_mag_asy'
THEMIS> tplot, ['wdc_mag_sym_0', 'wdc_mag_sym_1', 'wdc_mag_asy_0', 'wdc_mag_asy_1']
```

### How to Plot (SPEDAS-GUI):

Step 1: Start SPEDAS GUI Program.  
 Step 2: Choose [FILE] -> [Load Data].  
 Step 3: Choose [IUGONET] Tab.  
 Step 4: Uncheck 'Use Single Day'.  
 Step 5: Set Start Time: '2012-03-05 00:00:00' and Stop Time: '2012-03-12 00:00:00'.  
 Step 6: Choose Instrument Type: 'geomagnetic\_field\_index'.  
 Step 7: Choose Data Type: 'ASY\_index', Site or parameter(s)-1: 'WDC\_kyoto' and parameter(s)-2: 'asy', 'sym'.  
 Step 8: Push [->] button. (Please wait a few minutes).  
 Step 9: Push [Done] button.  
 Step 10: Choose [Graph] -> [Plot Layout Options].  
 Step 11: Choose 'wdc\_mag\_asy', 'wdc\_mag\_sym' and push [Line->] button.  
 Step 12: Push [OK] button.

4'. Advance (know how to analyze data by SPEDAS).

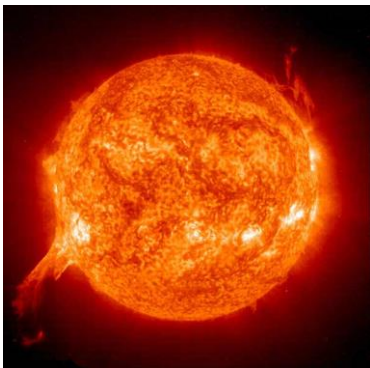
## Space Physics Environment Data Analysis Software (SPEDAS) - Page: 1



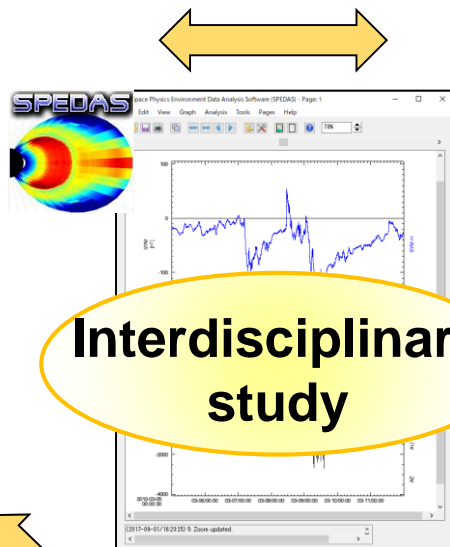
5. Analyze data by SPEDAS

- **SPEDAS (Space Physics Environment Data Analysis Software)** was developed by scientists and programmers of the UC Berkeley's Space Sciences Laboratory, UCLA's IGPP and other contributors.
- IUGONET has provided the plugin software for SPEDAS to handle the IUGONET data.
- It is useful for the interdisciplinary study about solar-terrestrial physics.

## Sun & Solar wind

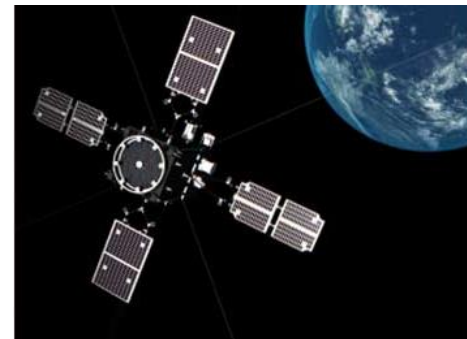


DSCOVR	SOHO	Wind	NASA OMNI
Stereo	ACE	IMP-8	....



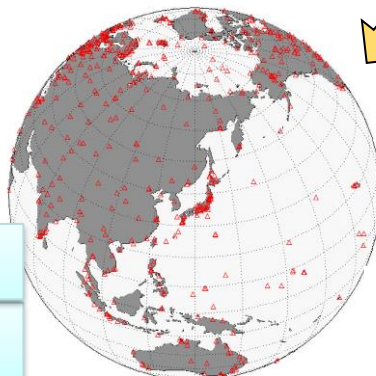
**Interdisciplinary study**

## Geospace



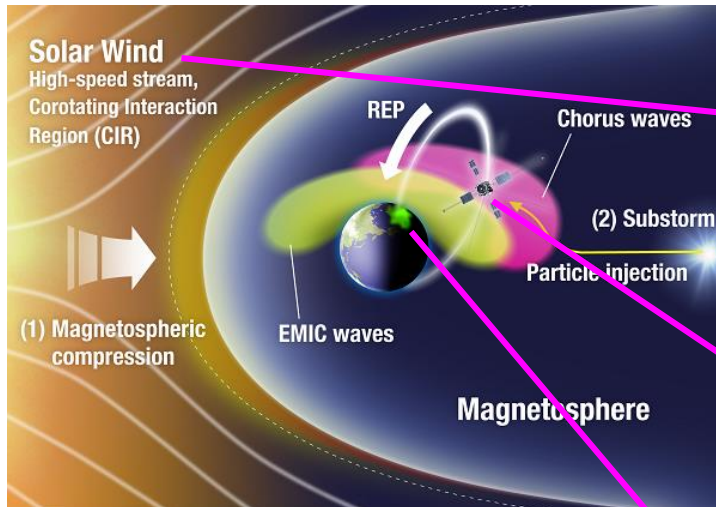
ERG	RBSP	THEMIS	LANL
GOES	MMS	POES	....

<b>IUGONET</b>	EISCAT	CARISMA	GIMA	ERG
Super DARN	THEMIS-GBO	MACCS	BARREL	....



**Earth's atmosphere**

Tanaka et al., 2019



Data from multiple regions

Data from multiple projects



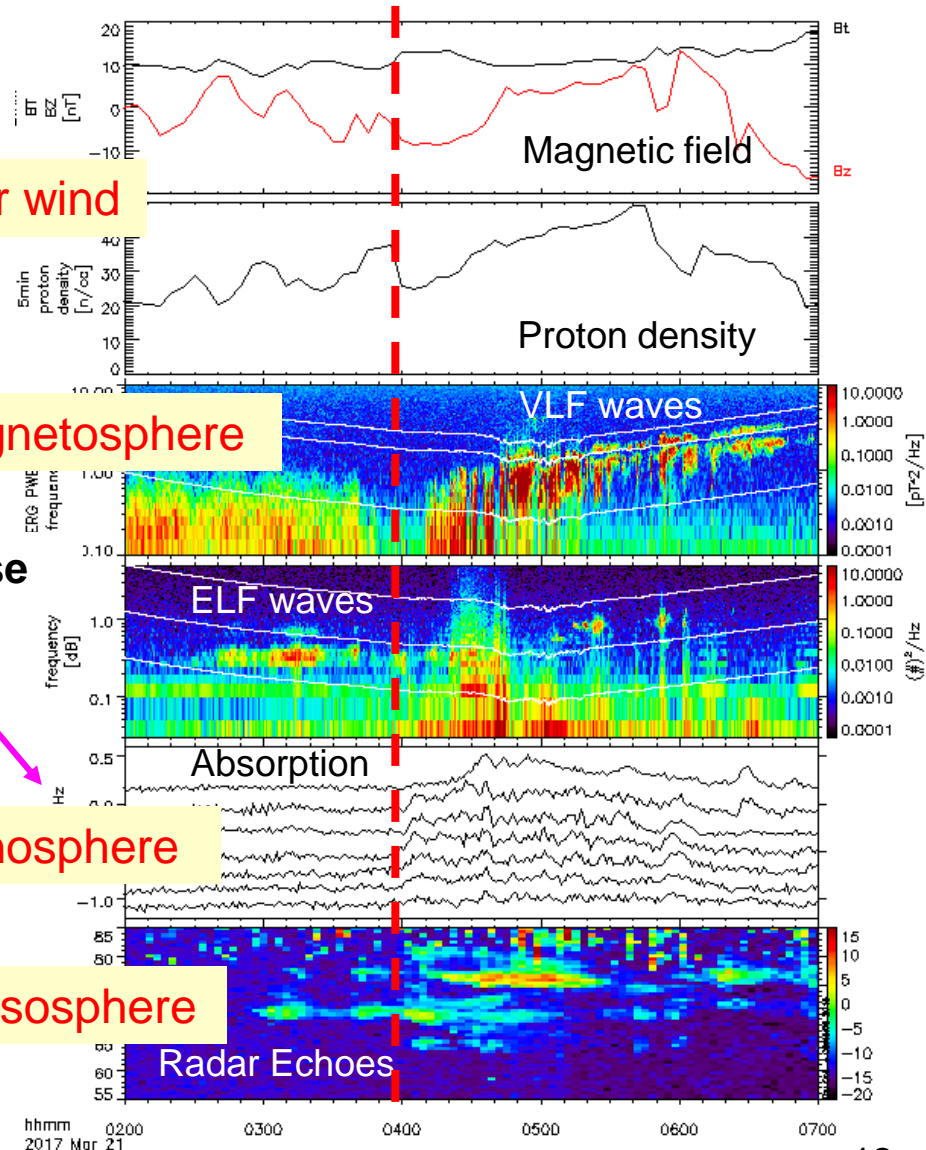
Solar wind

Magnetosphere

Arase

Ionosphere

Mesosphere

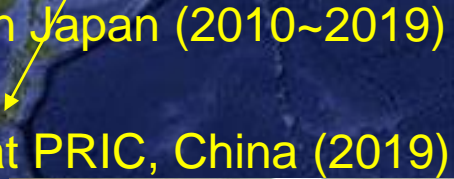
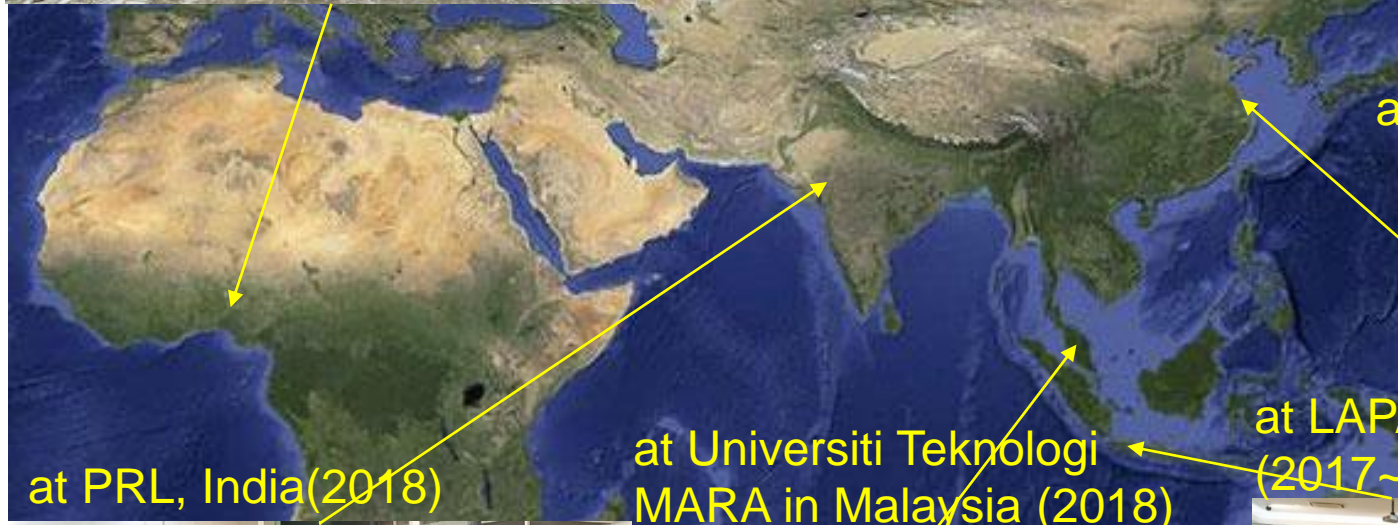




# Outreach activity



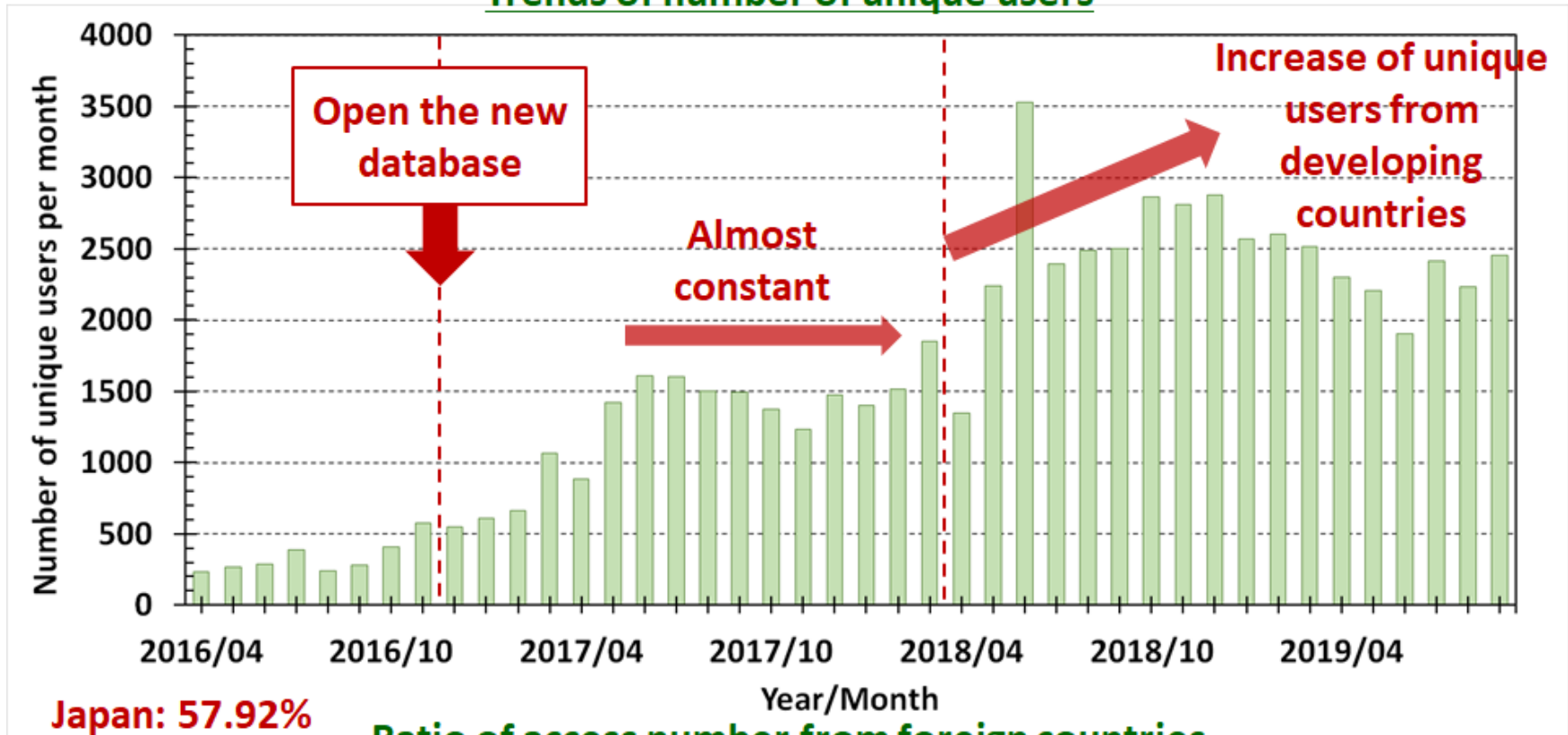
Education of young researchers and construction of international network



# Outcome (1)

## Access number of unique users to IUGONET Type-A

### Trends of number of unique users



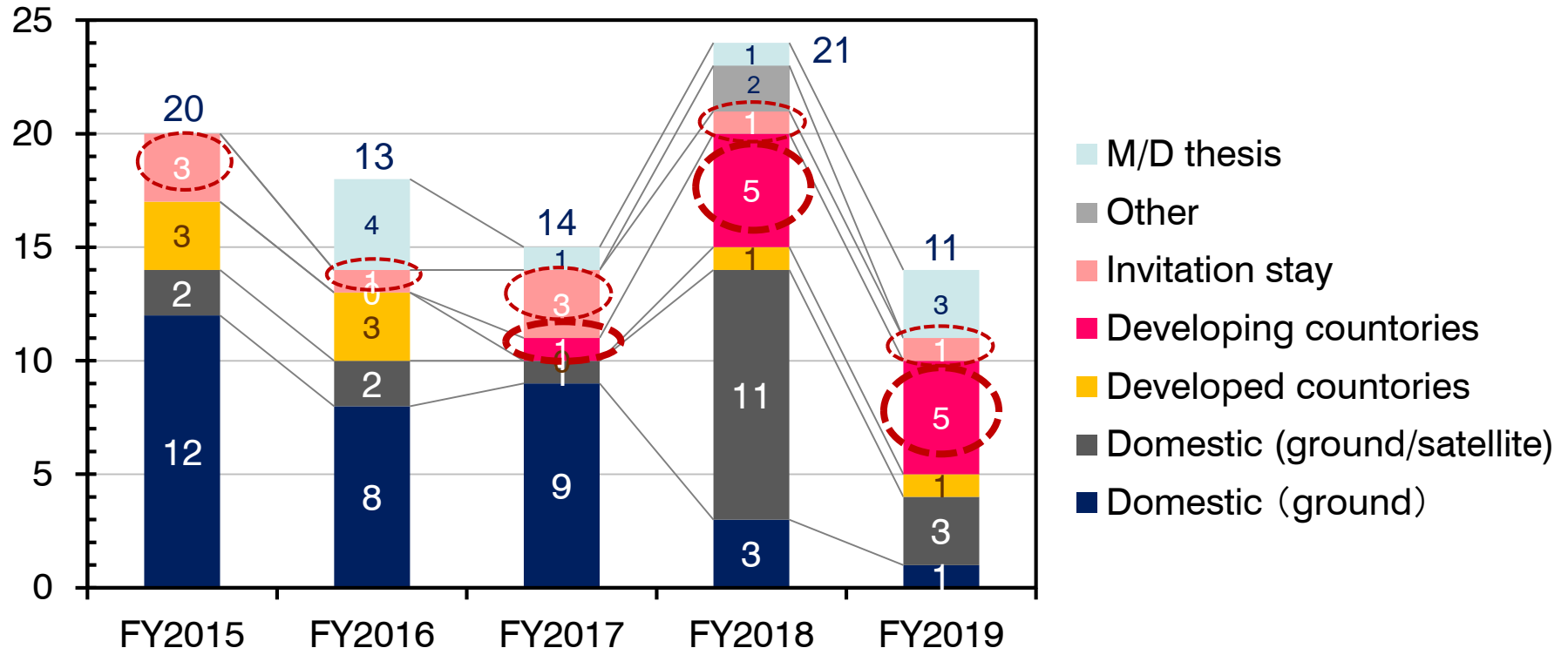
Japan: 57.92%  
Foreign: 42.08%

### Ratio of access number from foreign countries (2016/10 — 2019/10)

Units (%)

USA	Indone sia	China	Germa ny	Russia	Turkey	Korea	France	India	Brazil
35.5	9.8	9.0	5.3	4.7	4.5	3.8	2.9	2.1	1.6

## Number of papers with acknowledgements to IUGONET



## List of papers whose first author is a collaborator in Indonesia (2020)

- Koushik, N., K. K. Kumar, C. Vineeth, G. Ramkumar, and K. V. Subrahmanyam, Meteor Radar Observations of Lunar Semidiurnal Oscillations in the Mesosphere Lower Thermosphere over Low and Equatorial Latitudes and their variability during Sudden Stratospheric Warming Events, *J. Geophys. Res.*, 125, e2019JA027736, doi:10.1029/2019JA027736, 2020.
- Sridharan, S., Equatorial upper mesospheric mean winds and tidal response to strong El Niño and La Niña, *J. Atmos. Sol.-Terr. Phys.*, 202, doi: 10.1016/j.jastp.2020.105270, 2020.
- Koushik, N., K. Kishore Kumar, Geetha Ramkumar, K. V. Subrahmanyam, G. Kishore Kumar, W. K. Hocking, Maosheng He, and Ralph Latteck, Planetary waves in the mesosphere lower thermosphere during stratospheric sudden warming: observations using a network of meteor radars from high to equatorial latitudes, *Climate Dynamics*, doi:10.1007/s00382-020-05214-5, 2020.



# Future Plan

	FY2020	FY2021	FY2022	FY2023	FY2024
<b>Support for STP Research</b>	Addition of new functions to tools Collaboration with other projects (Master plan, Transformative Research Areas, etc.) Release of M-UDAS Enhancement of M-UDAS and SPEDAS-plugin IUGONET Type-A'				
<b>Promotion of data publication</b>	Addition of new observational data from other projects to Type-A Development of pipeline of data publication Start of operation of EISCAT_3D Campaign observations				
<b>Education of young researchers &amp; international collaboration</b>	Enhancement of international collaboration with universities/institutes in Asia, Oceania, and Africa Data analysis workshops International school				
<b>Timeline of other STP projects</b>	PWING Challenging Exploratory Research Projects for the Future (ROIS) SCOSTEP PRESTO Master Plan (EMU radar, EISCAT_3D, Ground-based observation networks) Transformative Research Areas				



- The IUGONET project has developed the advanced tools for upper atmospheric research, such as the metadata database (**IUGONET Type-A**) and the integrated data analysis software (**SPEDAS**).
- To achieve scientific results effectively, our strategy is to connect seamlessly the research procedures using IUGONET Type-A and SPEDAS.
- We have contributed to the education for young researchers both in Japan and other countries (e.g., Indonesia, Malaysia, India, China, Nigeria) to enhance the international collaborations in the STP field.
- As a result, the number of papers whose first author is a collaborator in the developing countries is increasing.
- We are planning to continue the support for the STP researches, the promotion of data publication, and the education of young researchers.