



ESA's Rosetta mission on the comet 67P/Churyumov-Gerasimenko

M. Fulchignoni



1986



1987



1991



Selected in 1993

1988



1989



Rosetta

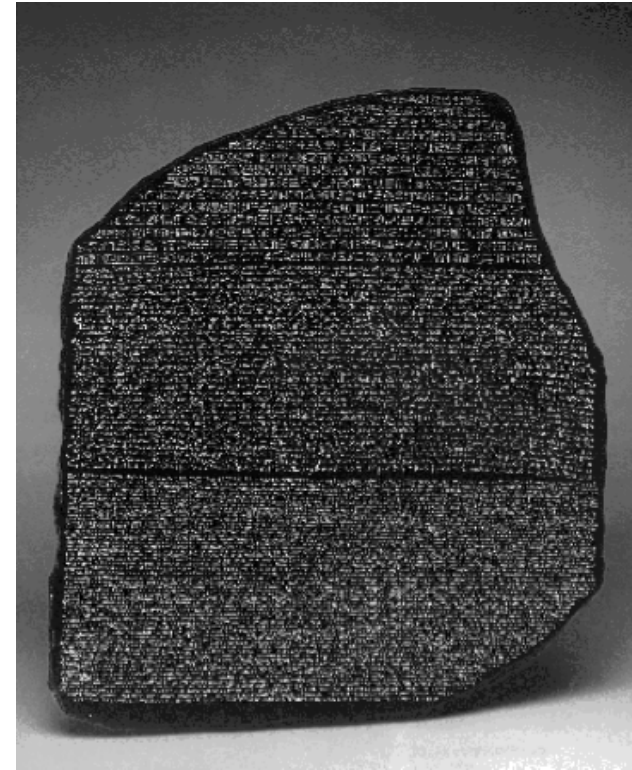
A comet rendezvous and two
asteroid flybys mission



Flying through and landing on a comet

ESA Rosetta mission

- First rendezvous to a comet
- Ambitious mission, ESA cornerstone aimed at the deciphering of our origins
- Scientific objectives
 - Characterize the evolution of activity from aphelion to perihelion
 - Determine the nucleus properties (structure, chemical & isotopic composition, physical & mineralogical properties), by orbiter payload & landing a modular laboratory on the still mysterious nucleus surface



Solar panels: 32m x 1m



Flying through and landing on a comet

ESA Rosetta mission

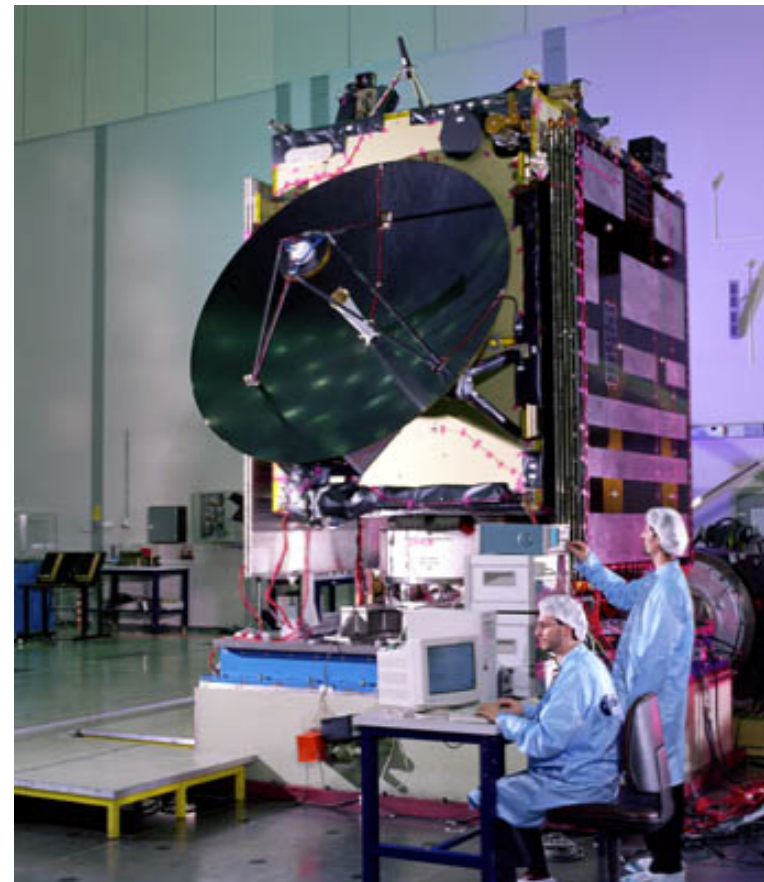
- Constraints related to very low gravity field of comet nuclei
 - ⇒ Relative velocity ≈ 0 at encounter
- Constraints on the trajectory
 - ⇒ Trajectory parallel to that of the nucleus or around it for ≈ 1 yrs
 - ⇒ Encounter near aphelion (smallest orbital velocity)
 - ⇒ Near ecliptic trajectory with perihelion near 1 AU (less demanding)
 - ⇒ Aphelion not too far away (Jupiter family comet)

Building Rosetta

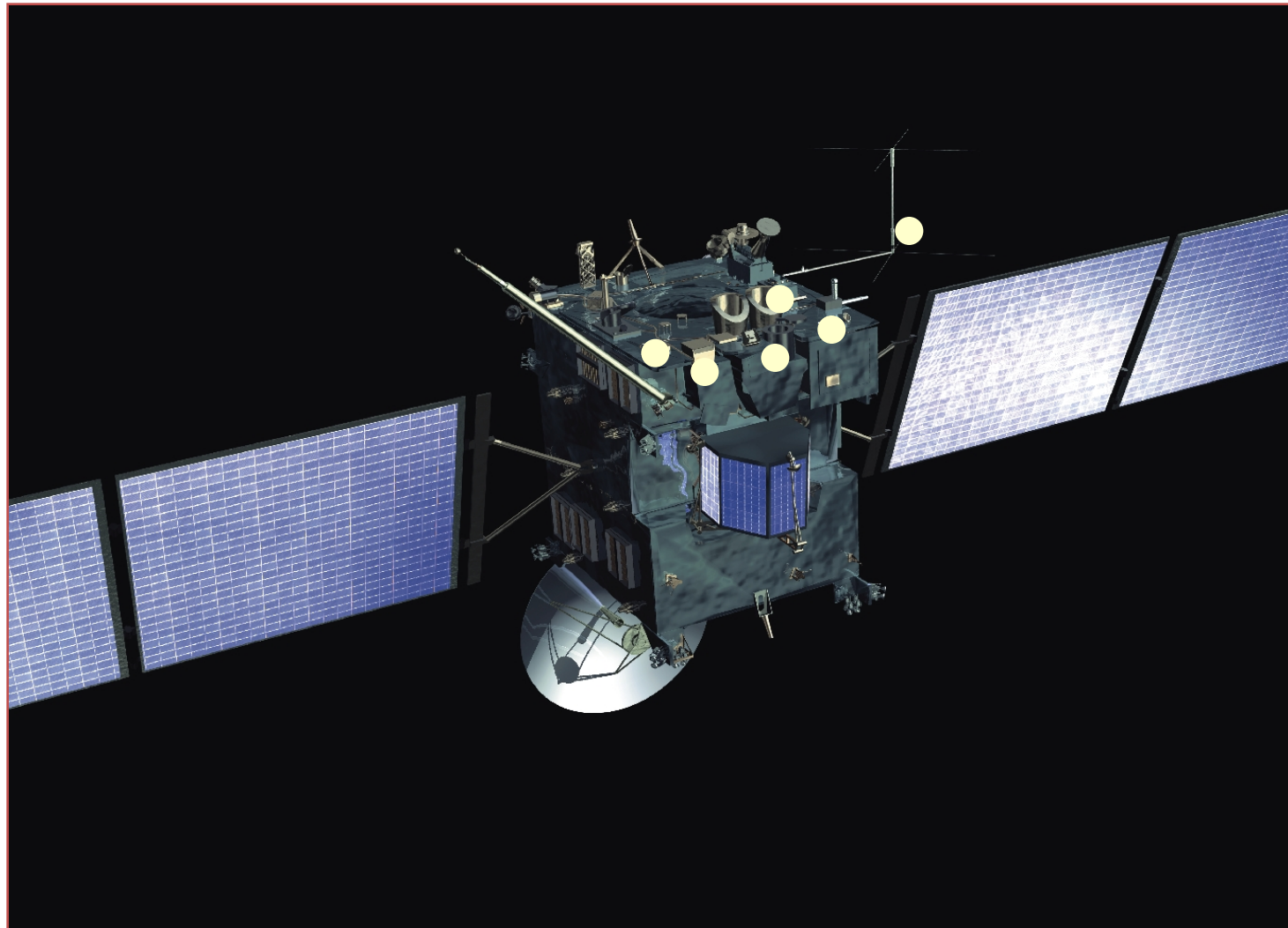
Massive and complex spacecraft

- Long duration mission
 - Long cruise phase (gravity assists and trajectory corrections)
 - Hibernation phase
 - Long rendezvous phase (2 years +)
 - About 10 sophisticated experiments on the spacecraft, as well as on the “Philae” lander
- ⇒ About 3 m x 2 m x 2 m
three axis stabilized spacecraft
⇒ 1.3 t (dry mass) + 1.6 t (fuel)
(Philae = 110 kg)

Integration the high gain antenna on a first model

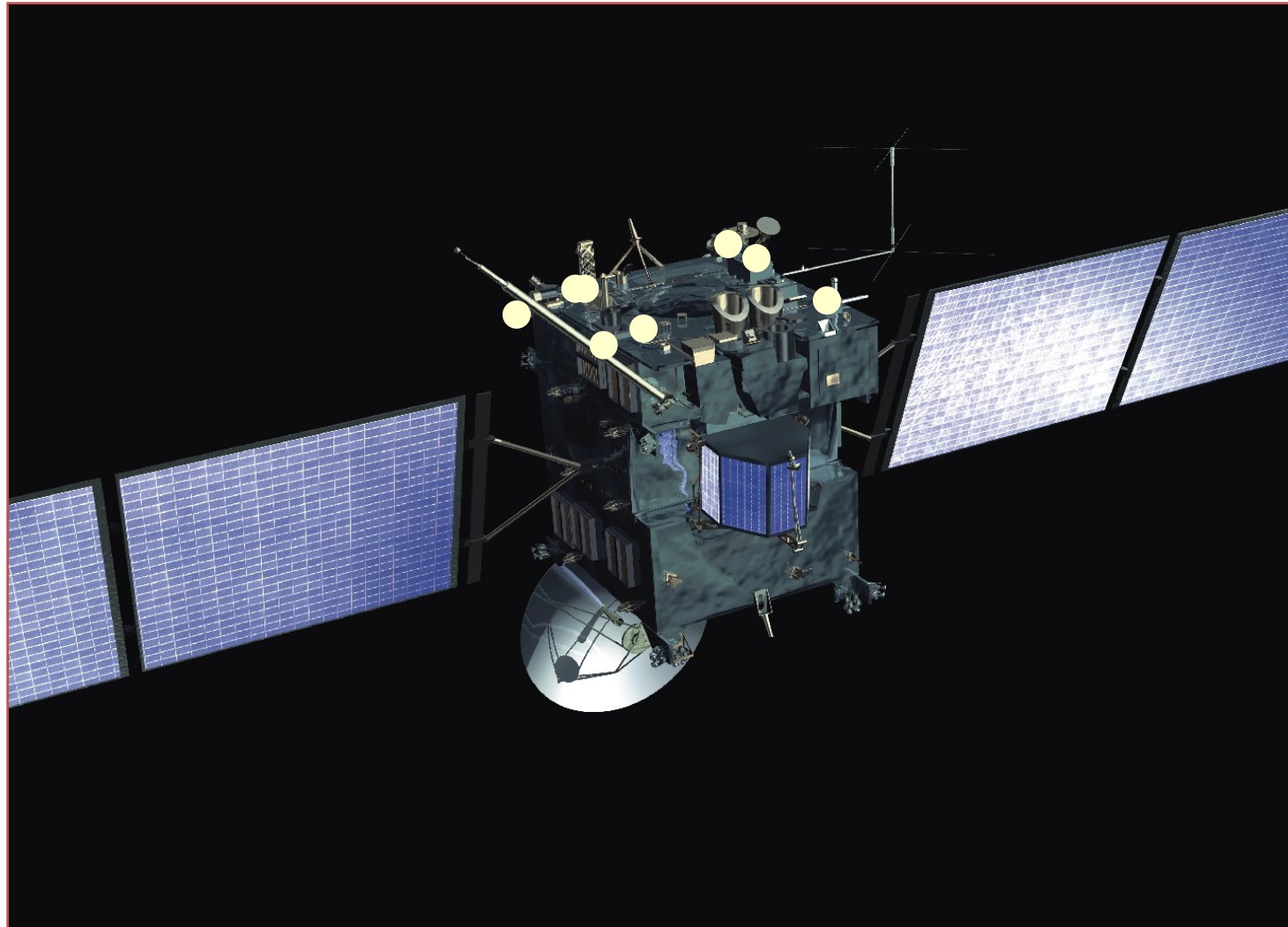


Rosetta Spacecraft and Payload



- OSIRIS: Optical, Spectroscopic, and Infrared Imaging Spectrograph (100 kg, 16 yr) many

Rosetta Spacecraft and Payload



● ROSINA (Rosetta Orbiter Spectrometer for Ion and Neutral Analysis) (ESA, Germany, Switzerland)

Philae Lander and Payload

Innovative small instruments for Philae

Imaging

Composition analysis

Physical properties

Nucleus large-scale structure

Magnetic field and plasma

Drill and sampling device

CIVA, ROLIS

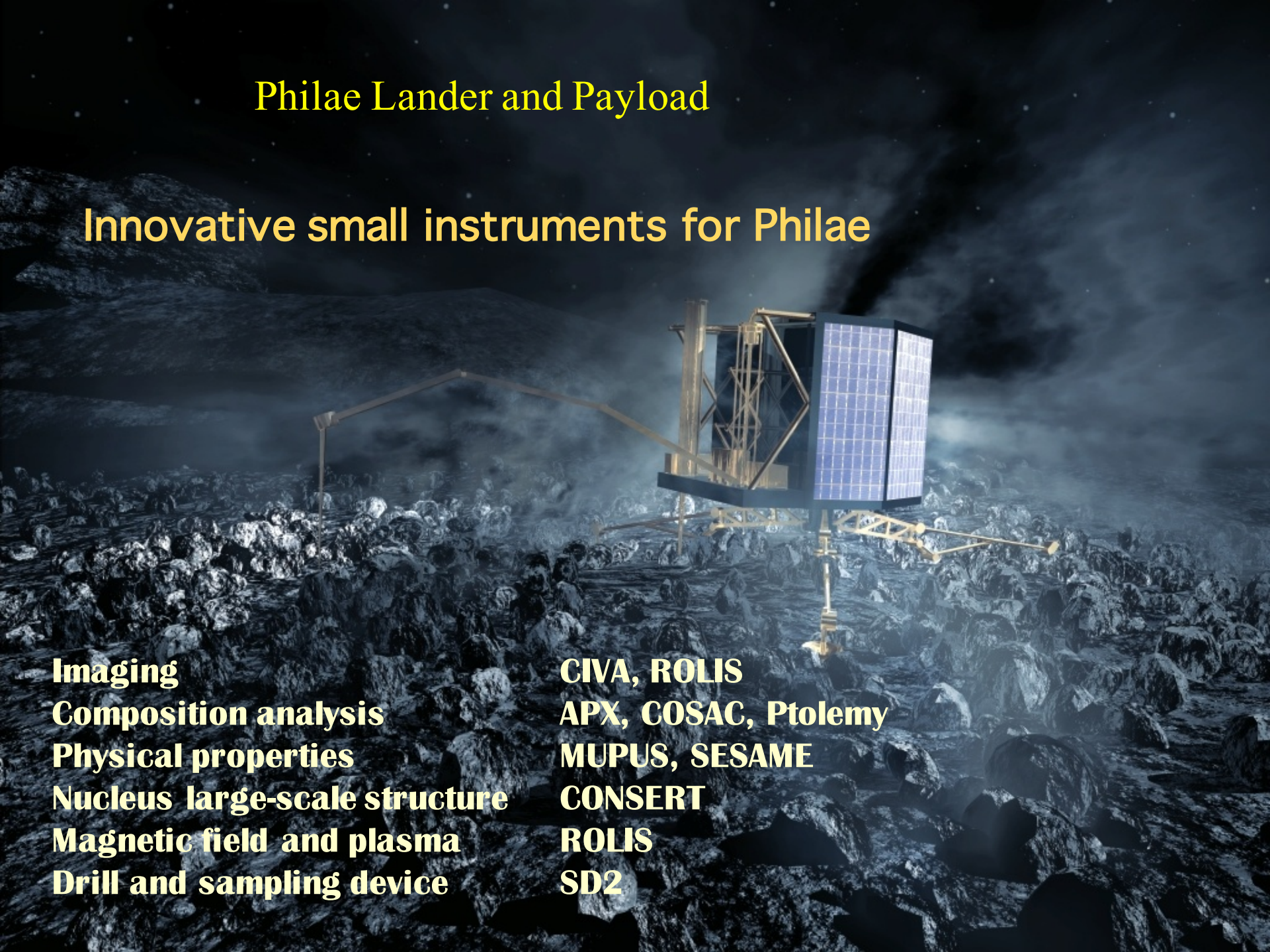
APX, COSAC, Ptolemy

MUPUS, SESAME

CONSERT

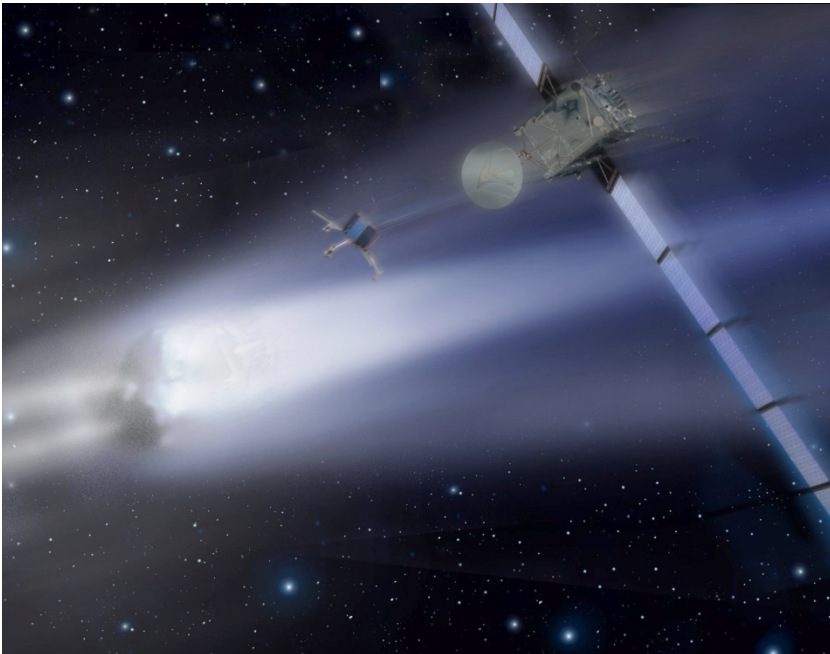
ROLIS

SD2

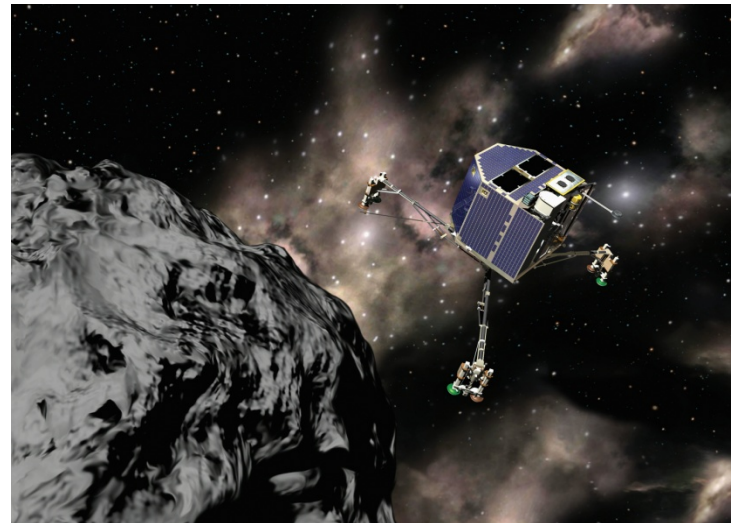


The Rosetta challenges

- Rendezvous with a low gravity object
- Landing on a mysterious object (unknown porosity, variable activity)
 - Survey from 50 km (surface mapping) down to 1 km
 - Descent of the probe and deployment of the landing gear with tripod, anchors and harpoons



ESA artist views of the mission



Temperature and solar energy constraints

- Sophisticated solar panels
32 m long x 1m large , 25 % efficiency



Journey to Churyumov-Gerasimenko

2003-2004, expecting the launch...

- Spacecraft fully integrated, flown to Kourou in September 2003
- Report of the launch to comet Wirtanen (anticipated to take place in January 2003) after an Ariane 5 failure



Journey to Churyumov-Gerasimenko

The new target

- Churyumov-Gerasimenko

Orbital parameters

$$a = 3.507 \text{ UA} \quad i = 7.12^\circ \quad e = 0.632$$

$$a^3/P^2 = 1 \text{ UA}^3 \text{ year}^{-2} \quad \Rightarrow P = 6.57 \text{ years}$$

$$p = a(1-e) \quad \Rightarrow p = 1.29 \text{ UA}$$

$$p' = a(1+e) \quad \Rightarrow p' = 5.74 \text{ UA}$$

- Physical properties
(estimated from HST observations)
Nucleus about 3 km x 5 km





Journey to Churyumov-Gerasimenko

Launch by Ariane 5G+

- One day report (meteorological and three more days report (insulation improvement))
- Successful launch, 2 March 2004
- After about 2 hours and 3 crossings of the radiation belts, successful upper stage ignition for interplanetary injection



Journey to Churyumov-Gerasimenko

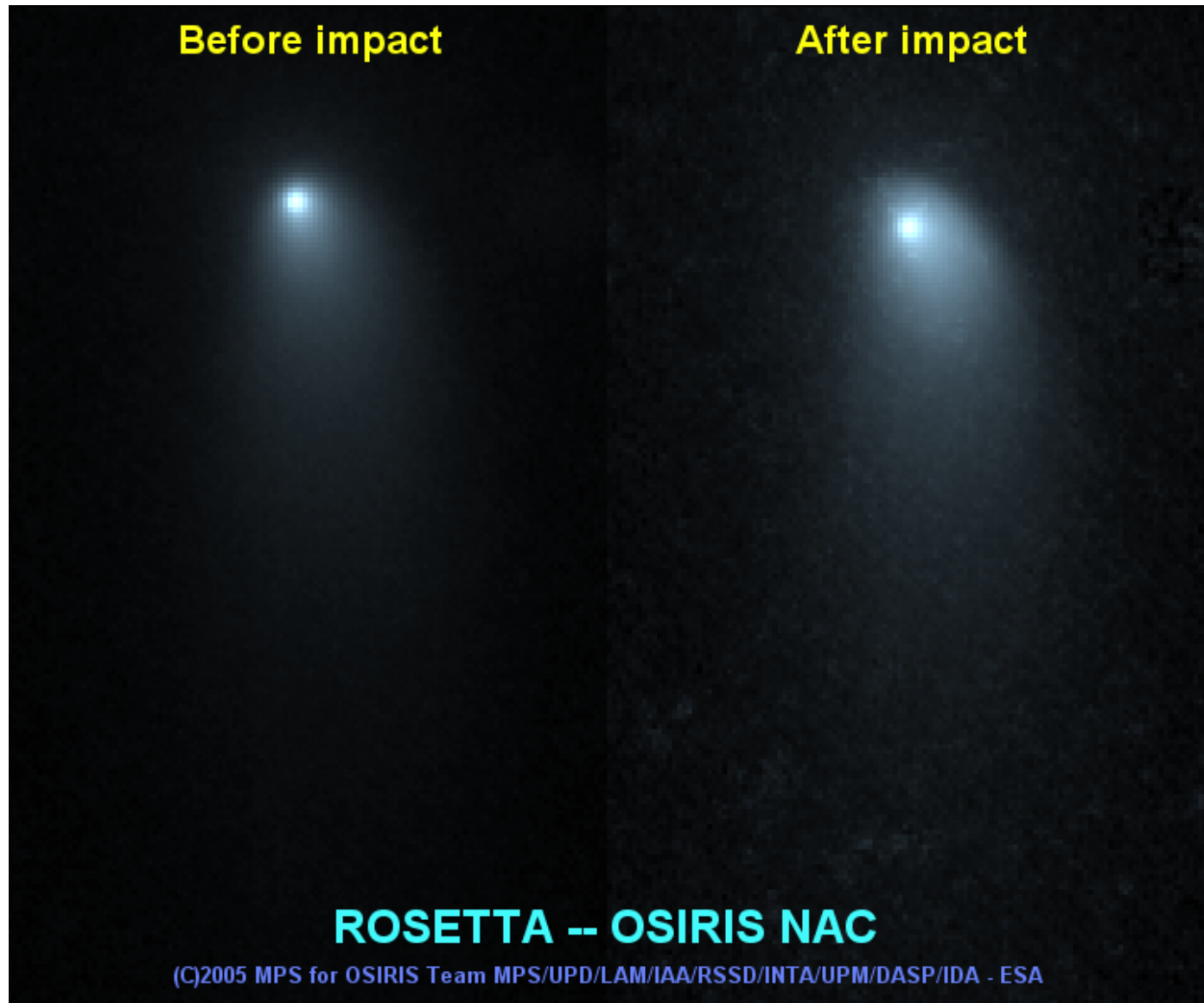
- All commissioning satisfactory
- Images from Osiris Camera of comet C/2002 T7 LINEAR



Rosetta was also seen from the Earth

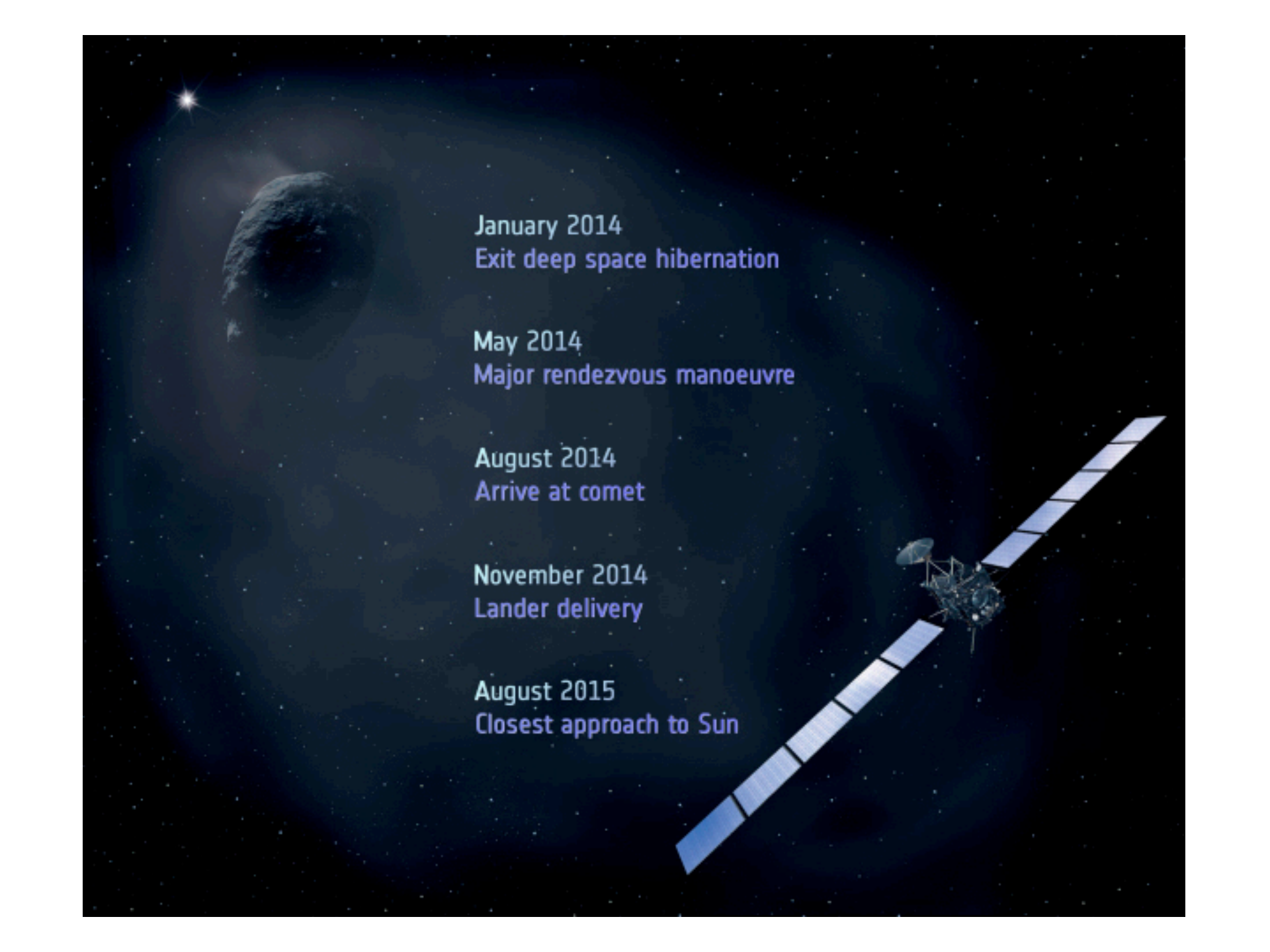


Deep impact observations in July 2005





***ROSETTA ON ITS WAY
TO
67P/CHURYUMOV-GERASIMENKO
IS VISITING
STEINS AND LUTETIA
ASTEROIDS***



January 2014
Exit deep space hibernation

May 2014
Major rendezvous manoeuvre

August 2014
Arrive at comet

November 2014
Lander delivery

August 2015
Closest approach to Sun





14 July 2014
Rot = 0 deg



5 km

→ COMET 67P/CHURYUMOV-GERASIMENKO'S VITAL STATISTICS

21.4 km³
Volume
 1.0×10^{13} kg
Mass
470 kg/m³
Density
70–80%
Porosity

2.6 km
2.3 km
1.8 km
1.8 km
3.3 km
4.1 km

4
Dust/gas ratio

5.3×10^{-4}
D/H ratio

Average water vapour production
300 ml/s → June 2014
600 ml/s → July 2014
1200 ml/s → August 2014

Rotation period
12.4043 hours

Spin axis:
69.3°
Right Ascension

64.1°
Declination

52°
Obliquity of the comet's rotational axis

X, Y Equatorial axes
Z Spin axis

–93°C to –43°C
Surface temperature

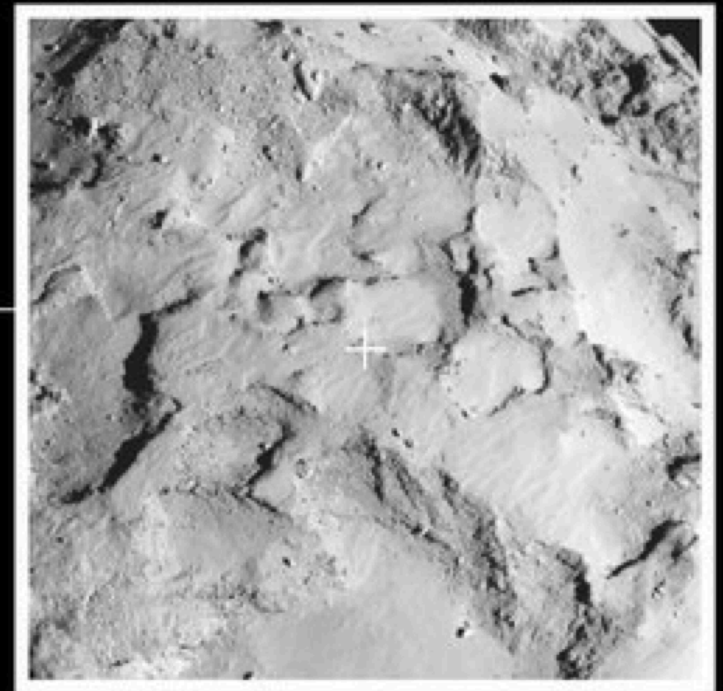
–243°C to –113°C
Subsurface temperature

6%
Average albedo

Rotation shape model: OSIRIS; surface temperature: VIRTIS; subsurface temperature: MIRO; water production rate: MIRO; D/H: ROSINA; dust/gas: GIADA; MIRO; ROSINA; volume: OSIRIS; mass: RSI; density: RSI/OSIRIS; albedo: OSIRIS; VIRTIS; comet images: NavCam

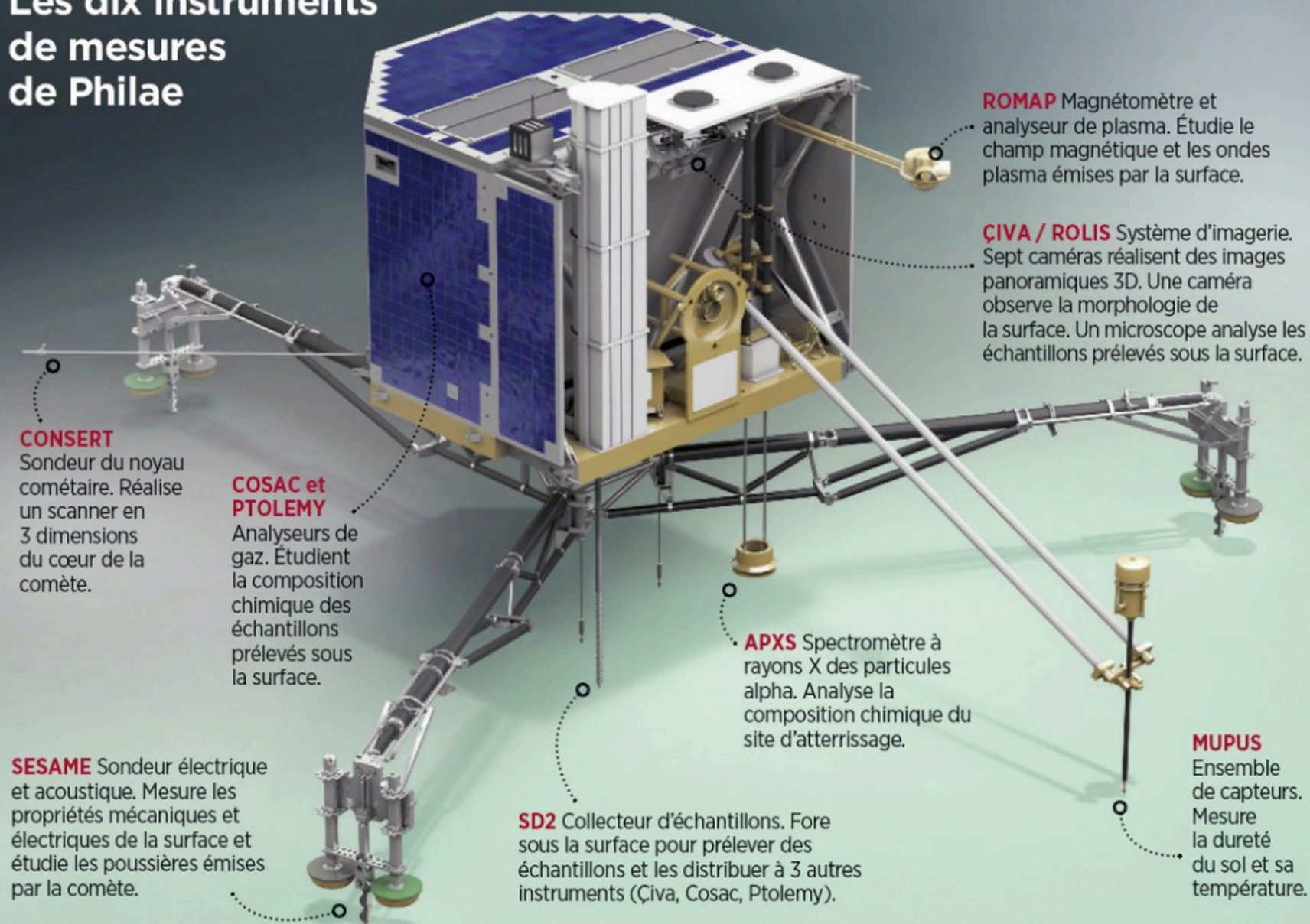
Data based on values published in January 2015

→ PHILAE'S LANDING SITE



Agilkia region

Les dix instruments de mesures de Philae



CONCERT

Sondeur du noyau cométaire. Réalise un scanner en 3 dimensions du cœur de la comète.

COSAC et PTOLEMY

Analyseurs de gaz. Étudient la composition chimique des échantillons prélevés sous la surface.

SESAME Sondeur électrique et acoustique. Mesure les propriétés mécaniques et électriques de la surface et étudie les poussières émises par la comète.

ROMAP Magnétomètre et analyseur de plasma. Étudie le champ magnétique et les ondes plasma émises par la surface.

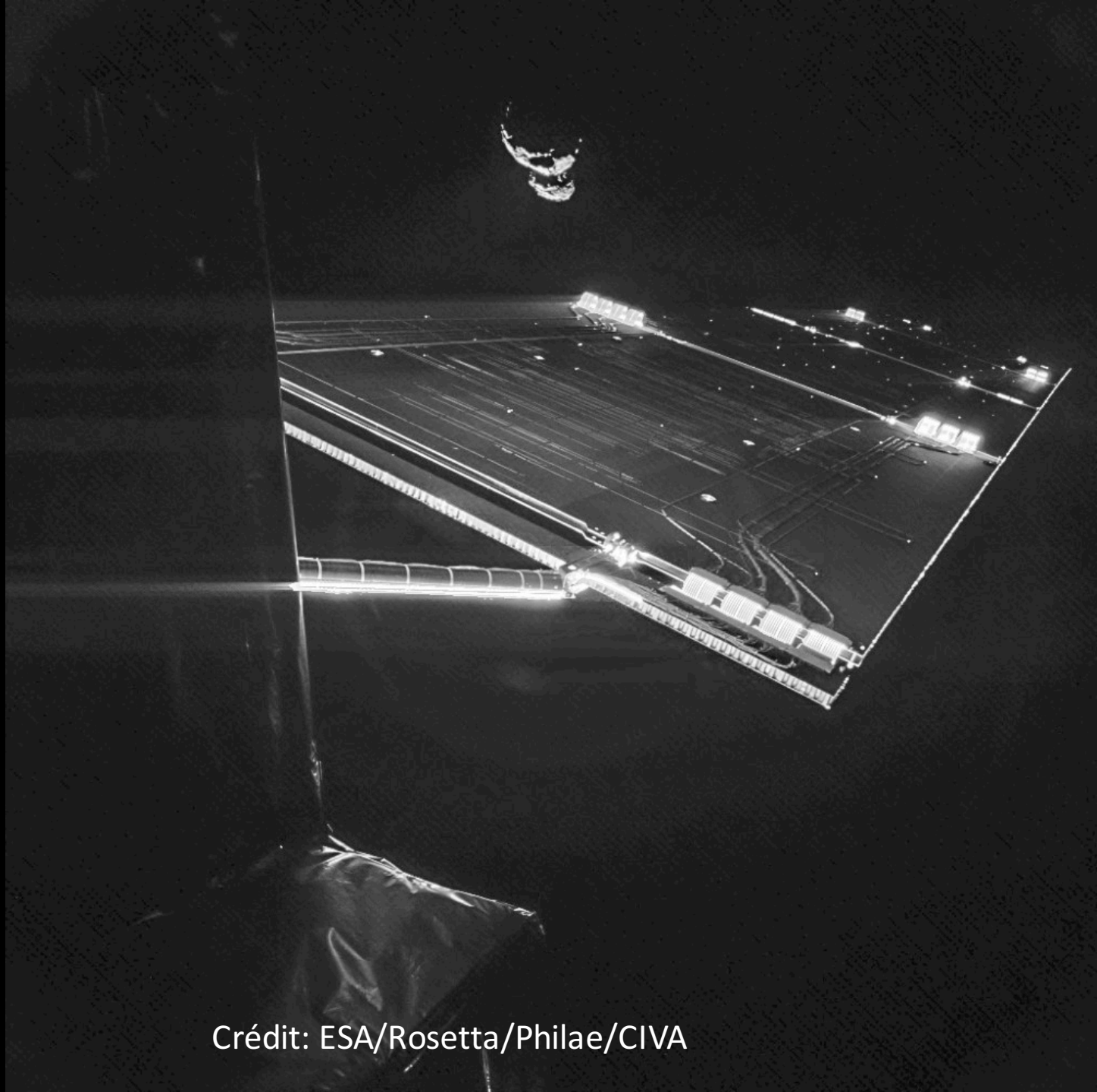
ÇIVA / ROLIS Système d'imagerie. Sept caméras réalisent des images panoramiques 3D. Une caméra observe la morphologie de la surface. Un microscope analyse les échantillons prélevés sous la surface.

APXS Spectromètre à rayons X des particules alpha. Analyse la composition chimique du site d'atterrissage.

SD2 Collecteur d'échantillons. Fore sous la surface pour prélever des échantillons et les distribuer à 3 autres instruments (Çiva, Cosac, Ptolemy).

MUPUS

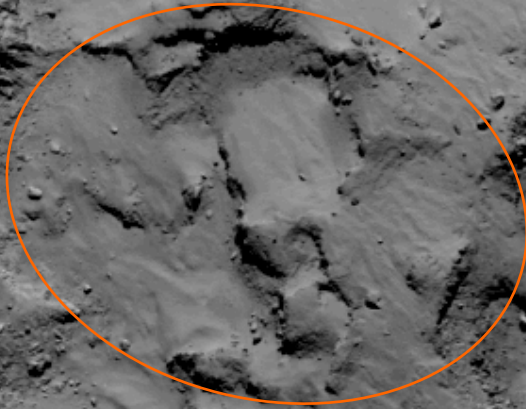
Ensemble de capteurs. Mesure la dureté du sol et sa température.



Crédit: ESA/Rosetta/Philae/CIVA



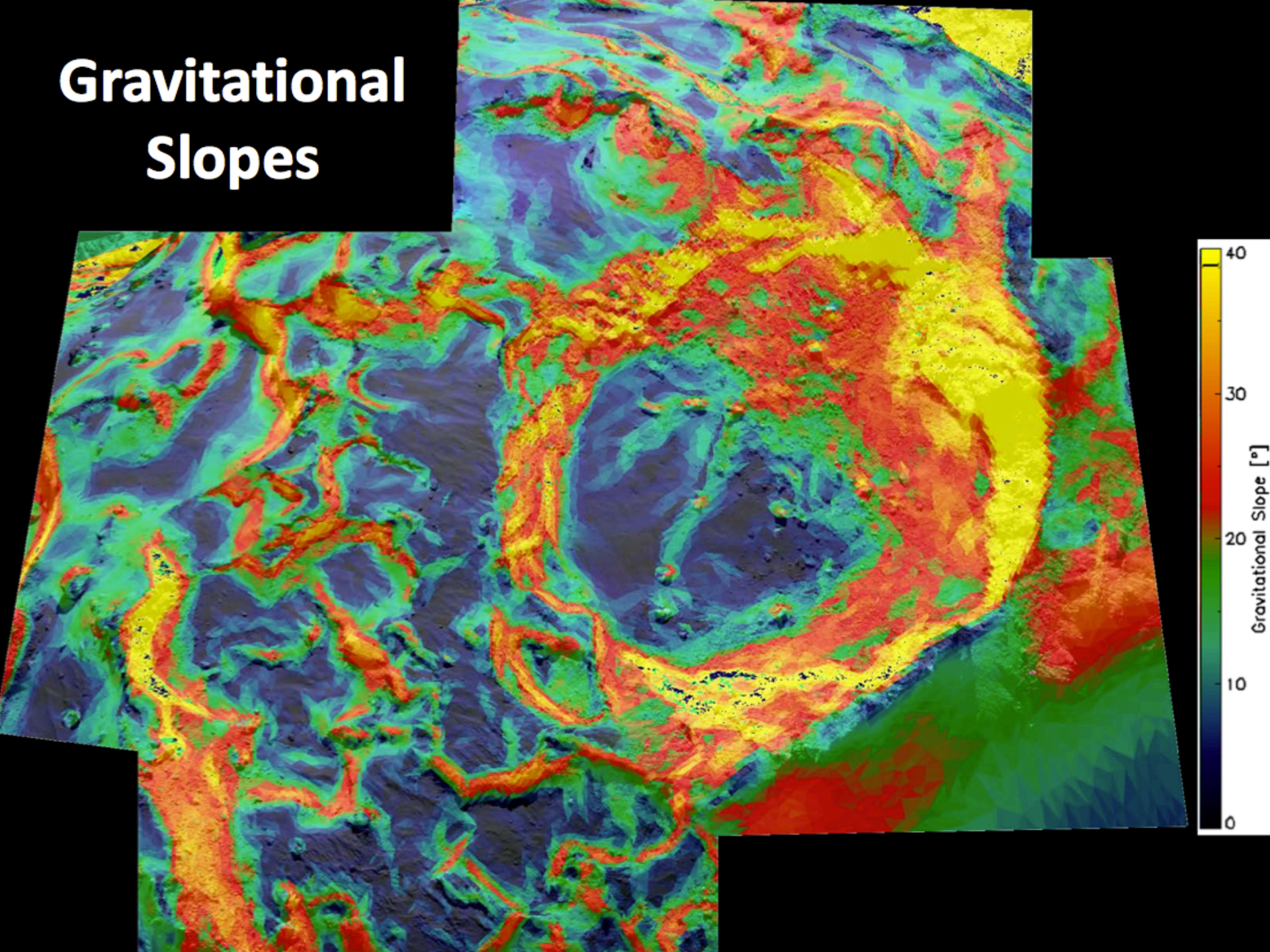
Agilkia region



100 m



Gravitational Slopes



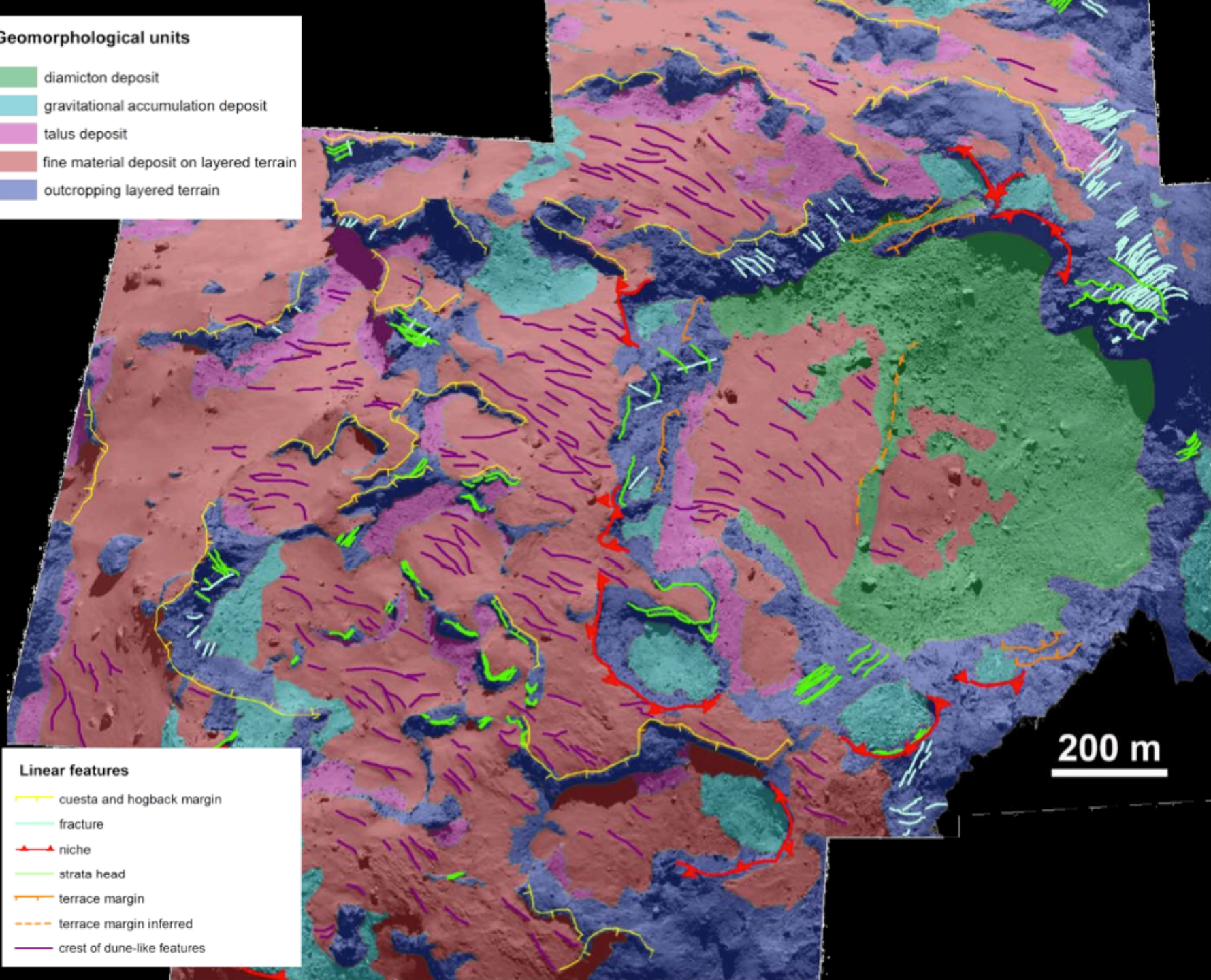
Geomorphological units

- diamicton deposit
- gravitational accumulation deposit
- talus deposit
- fine material deposit on layered terrain
- outcropping layered terrain

Linear features

- cuesta and hogback margin
- fracture
- niche
- strata head
- terrace margin
- terrace margin inferred
- crest of dune-like features

200 m







NAVCAM
15:30-15:34

1. 250m

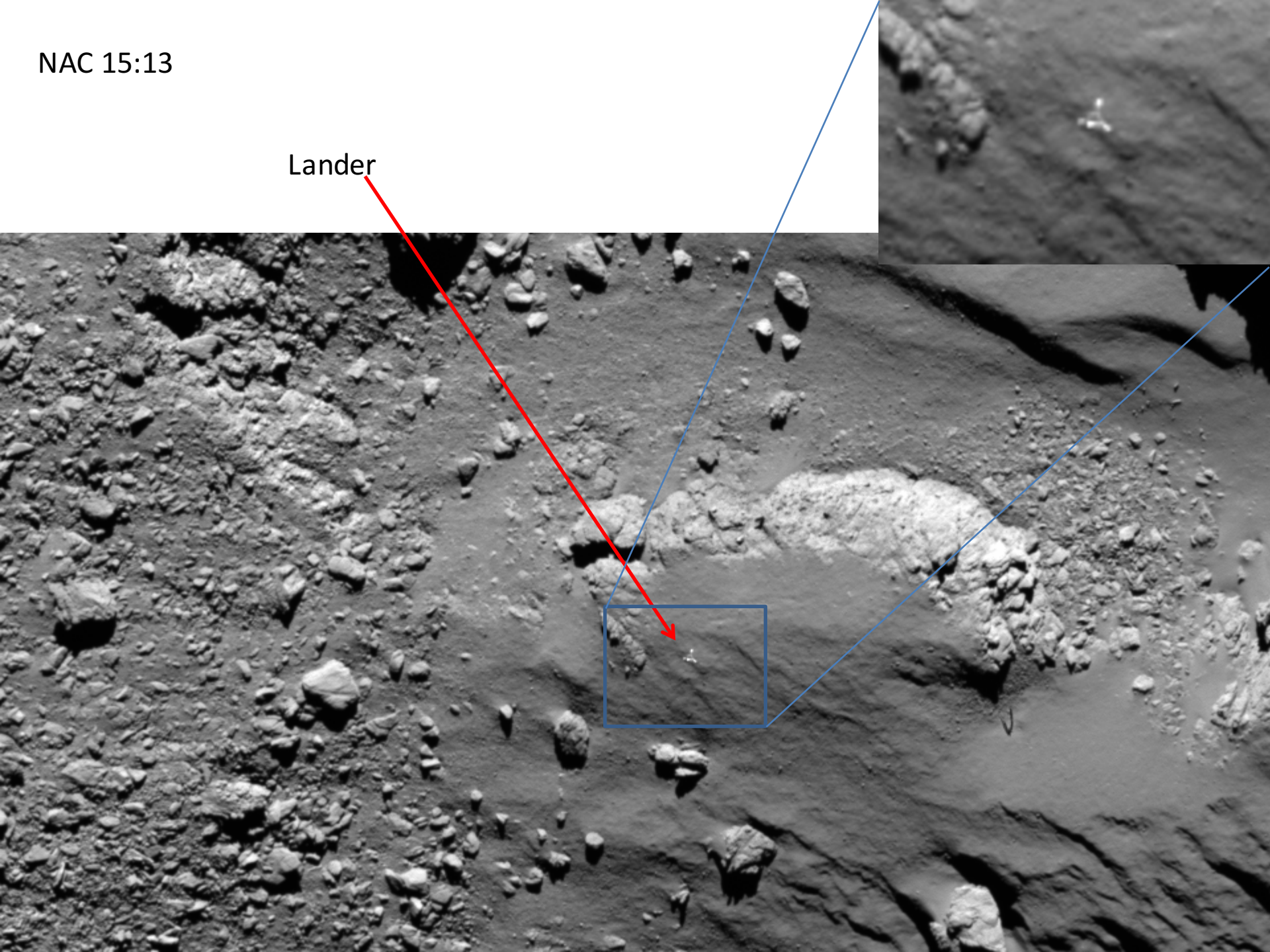
Rosetta @15km

Res=1.5 m

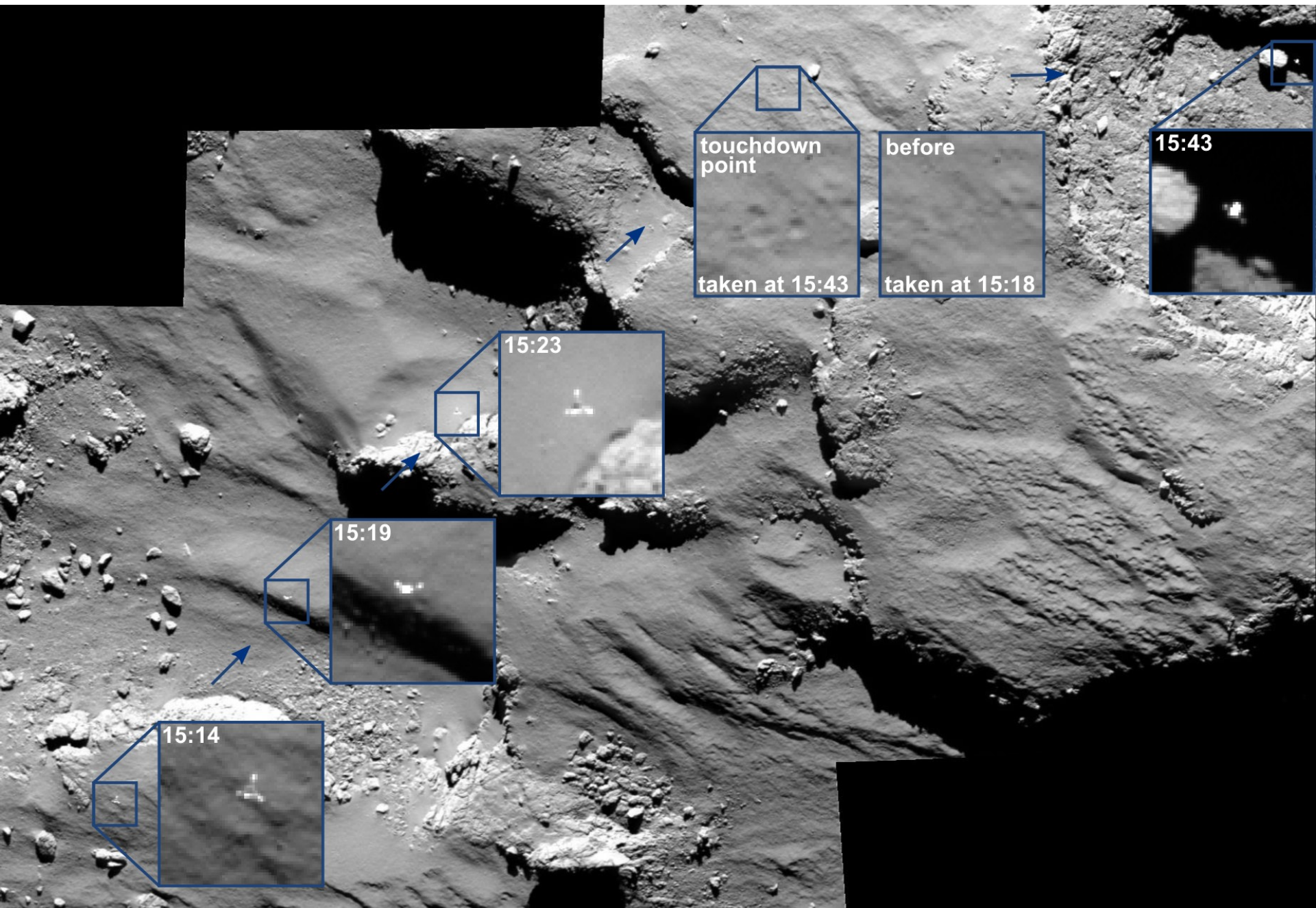


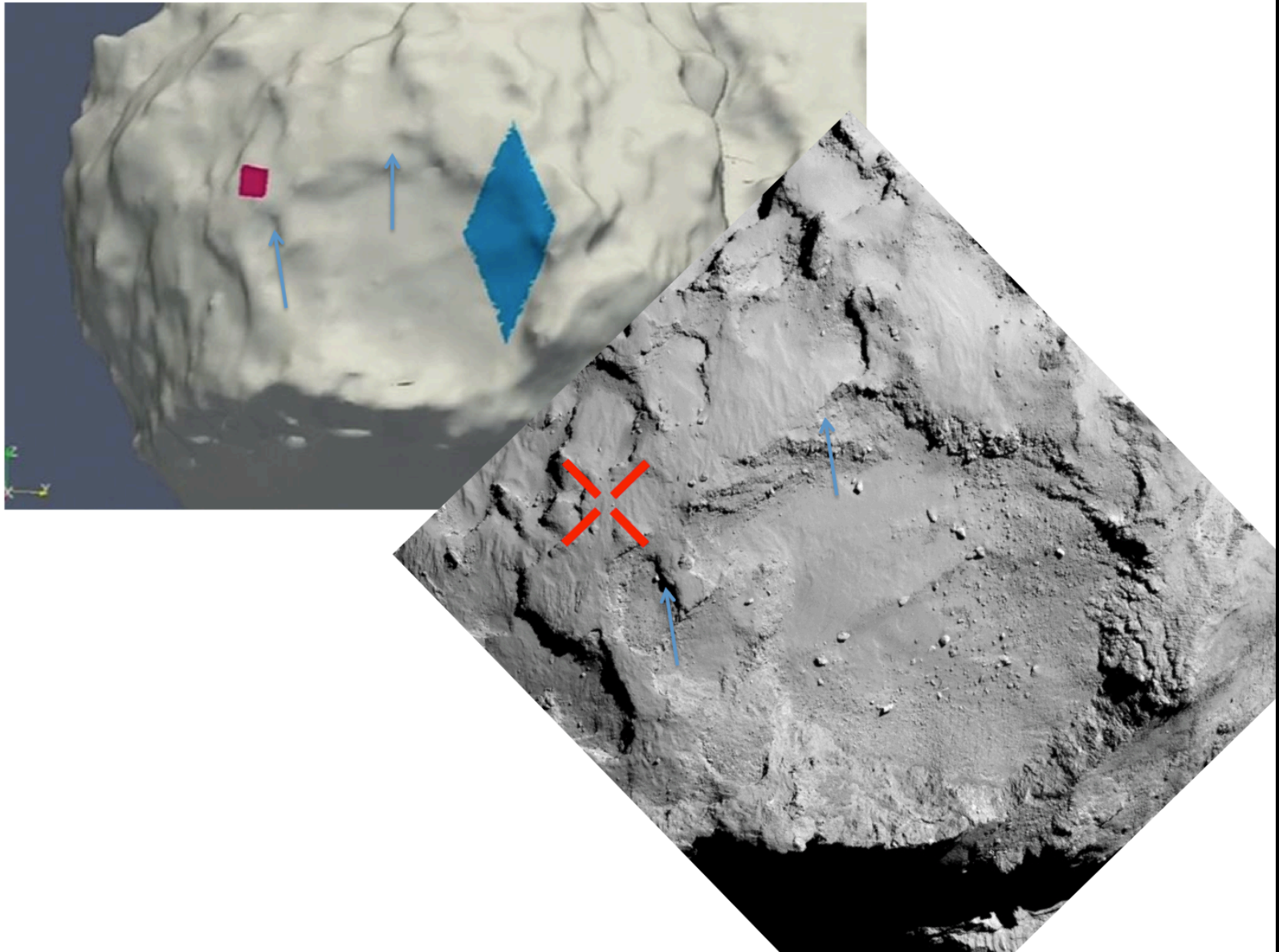
NAC 15:13

Lander

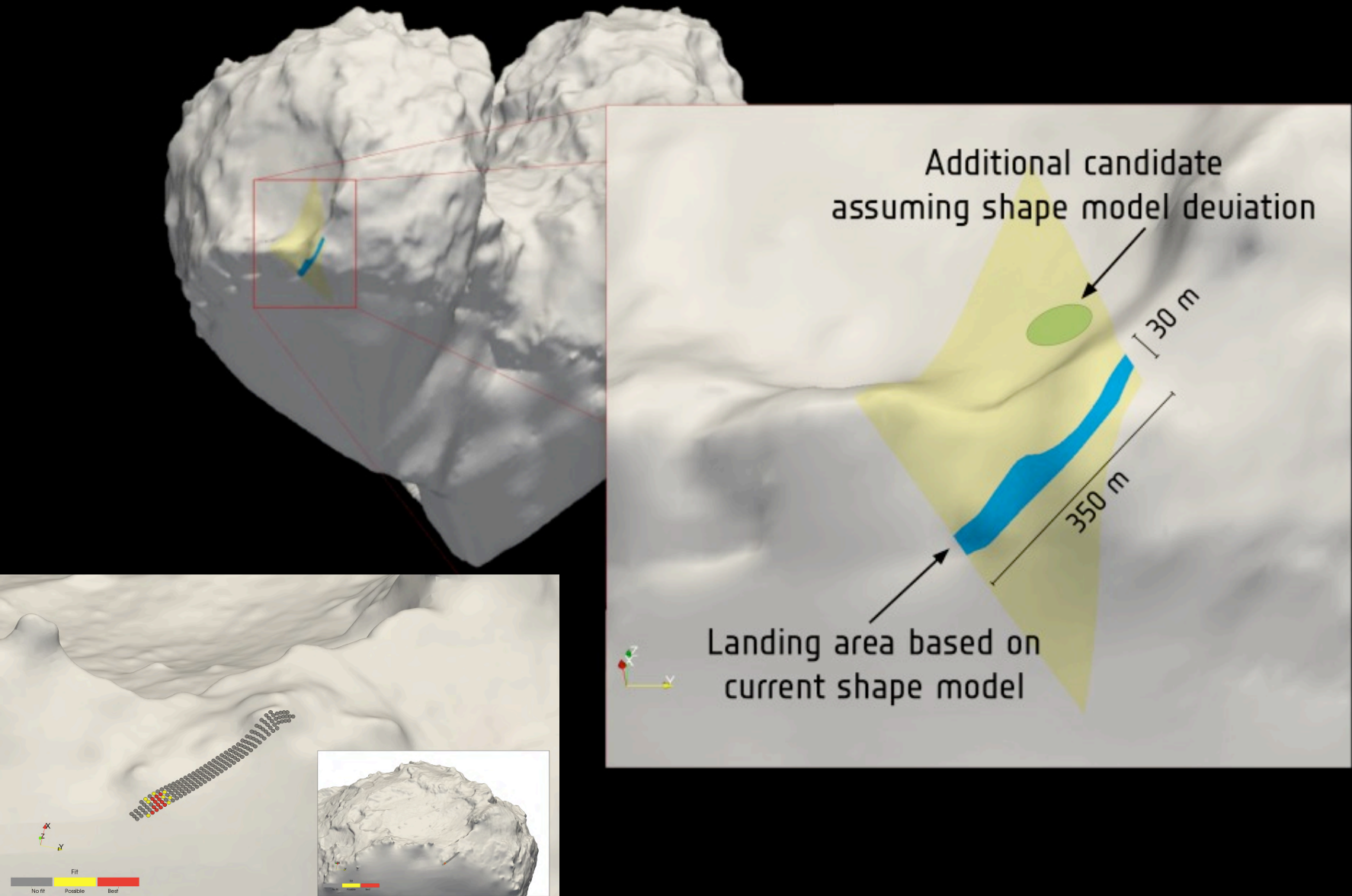


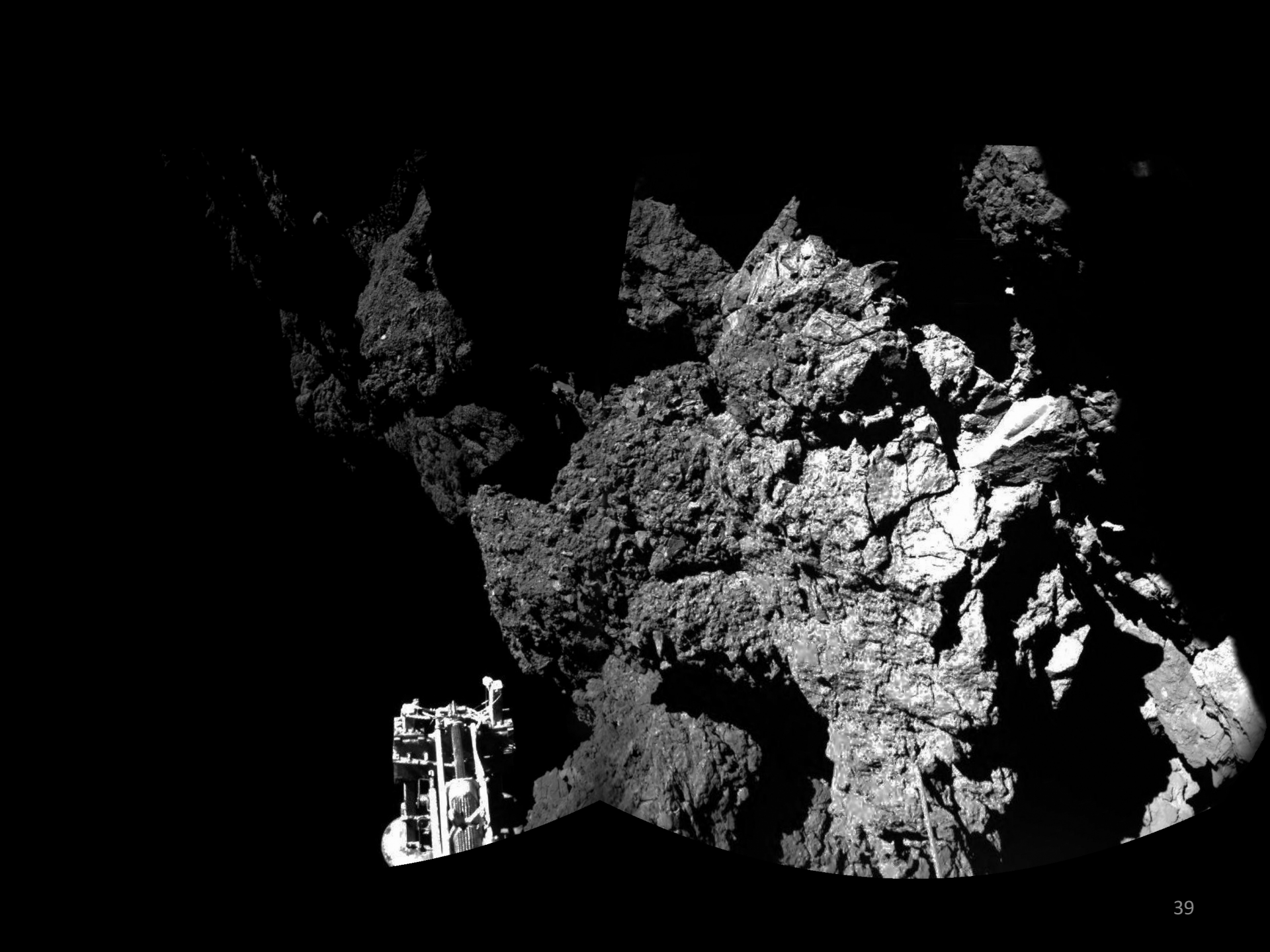
Touchdown 15:34

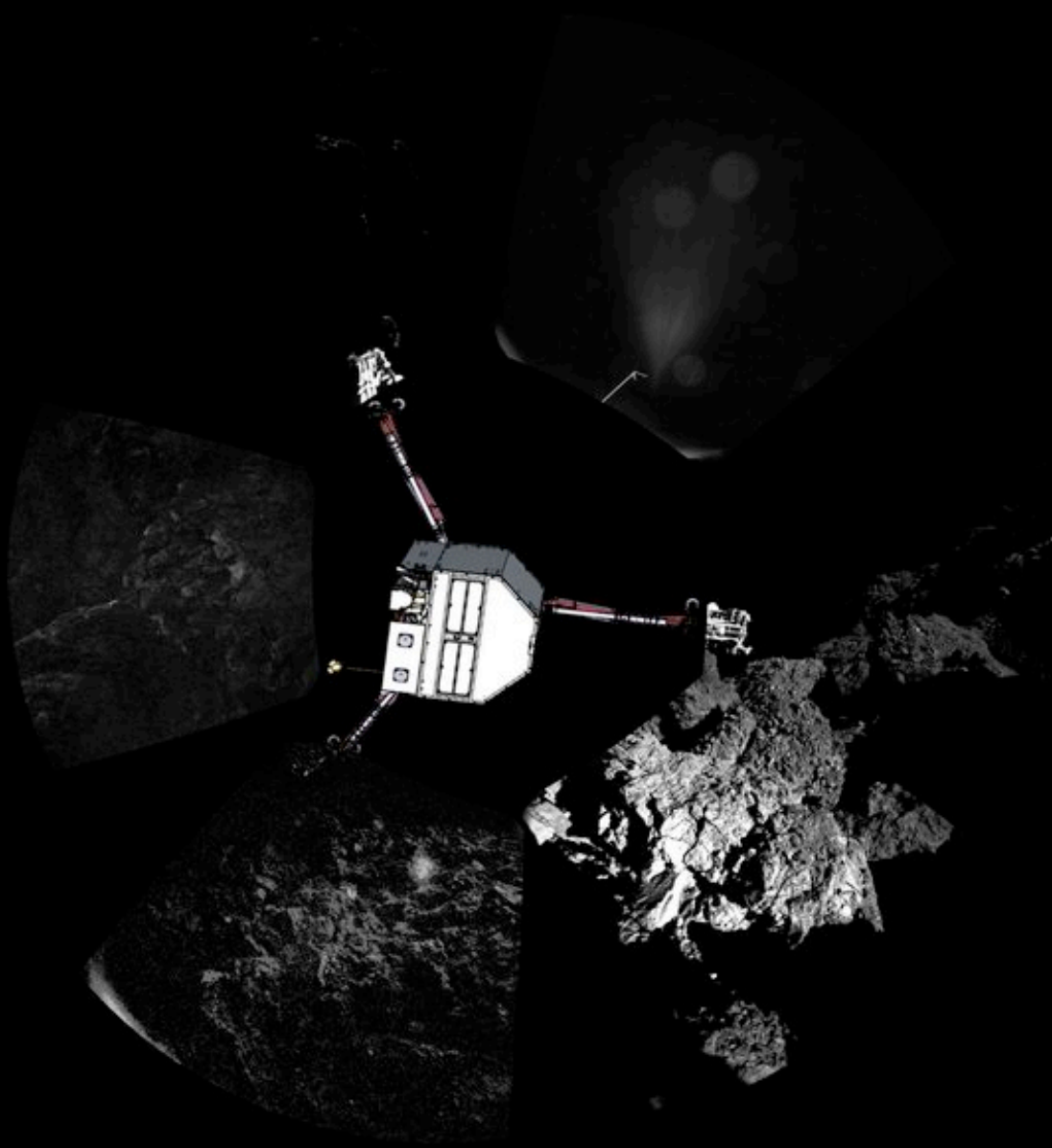




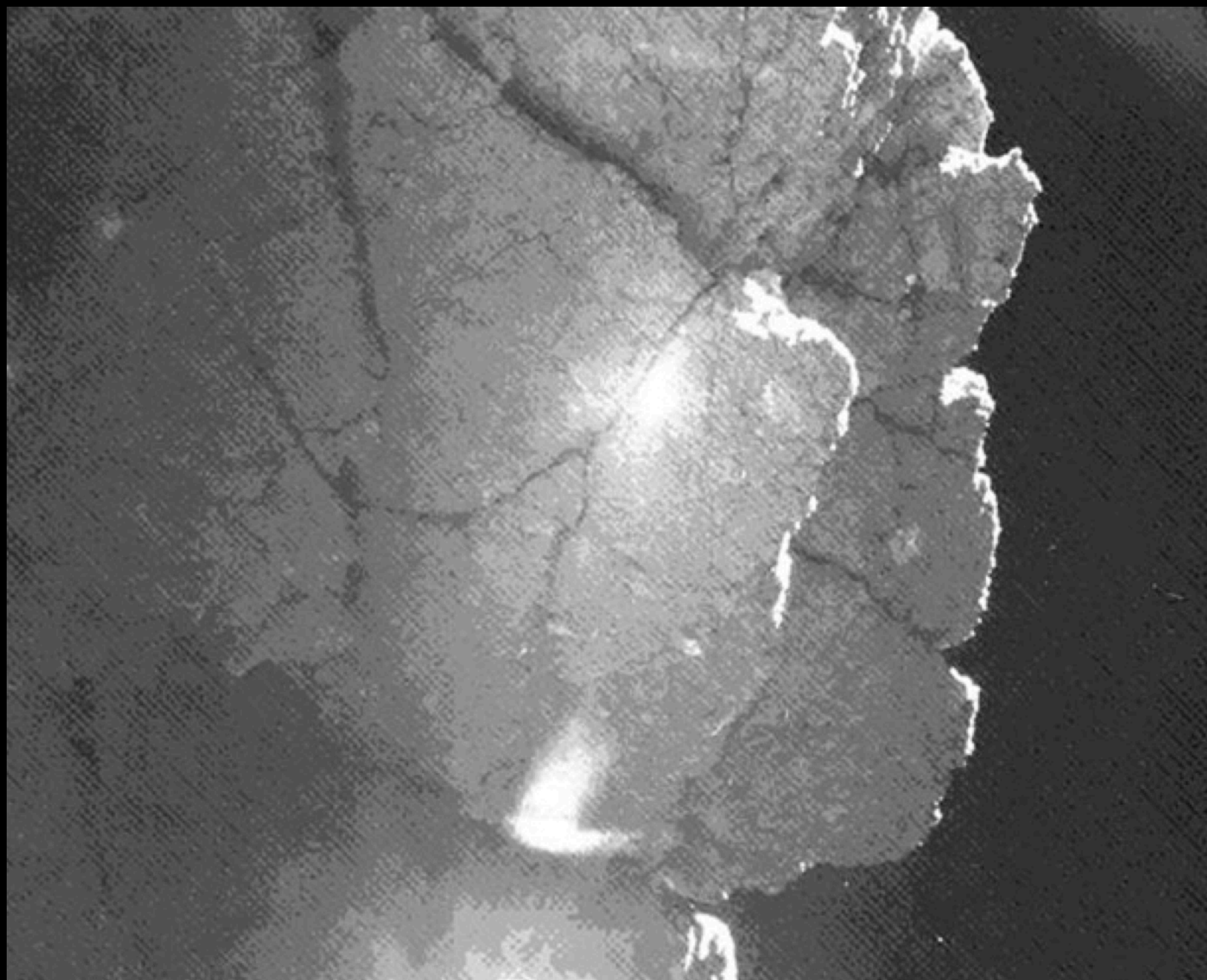
→ CONCERT estimation of landing area

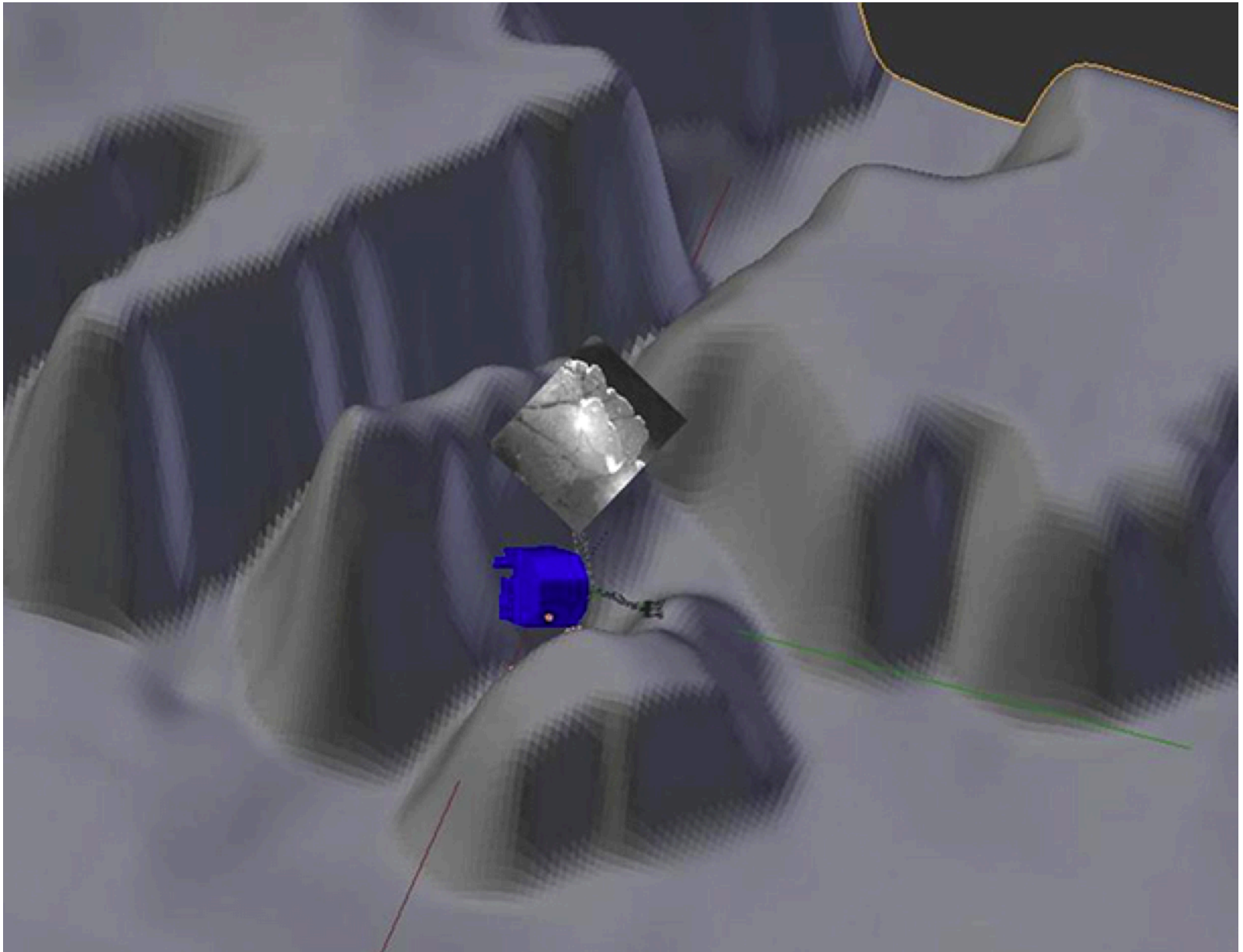






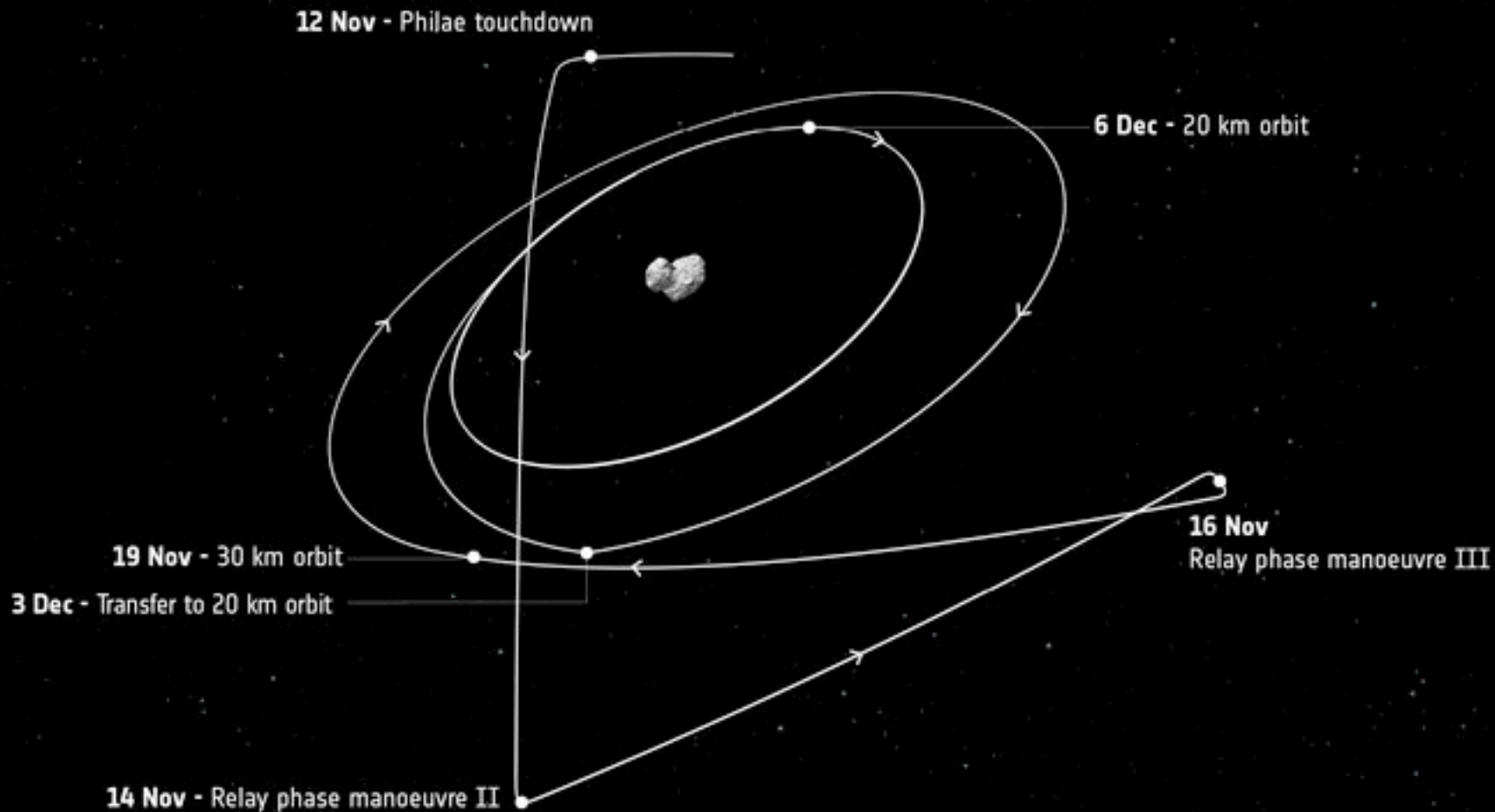
Superimposed on top of the image is a sketch of the Philae lander in the configuration the lander team currently believe it is in.





CNES reconstruction



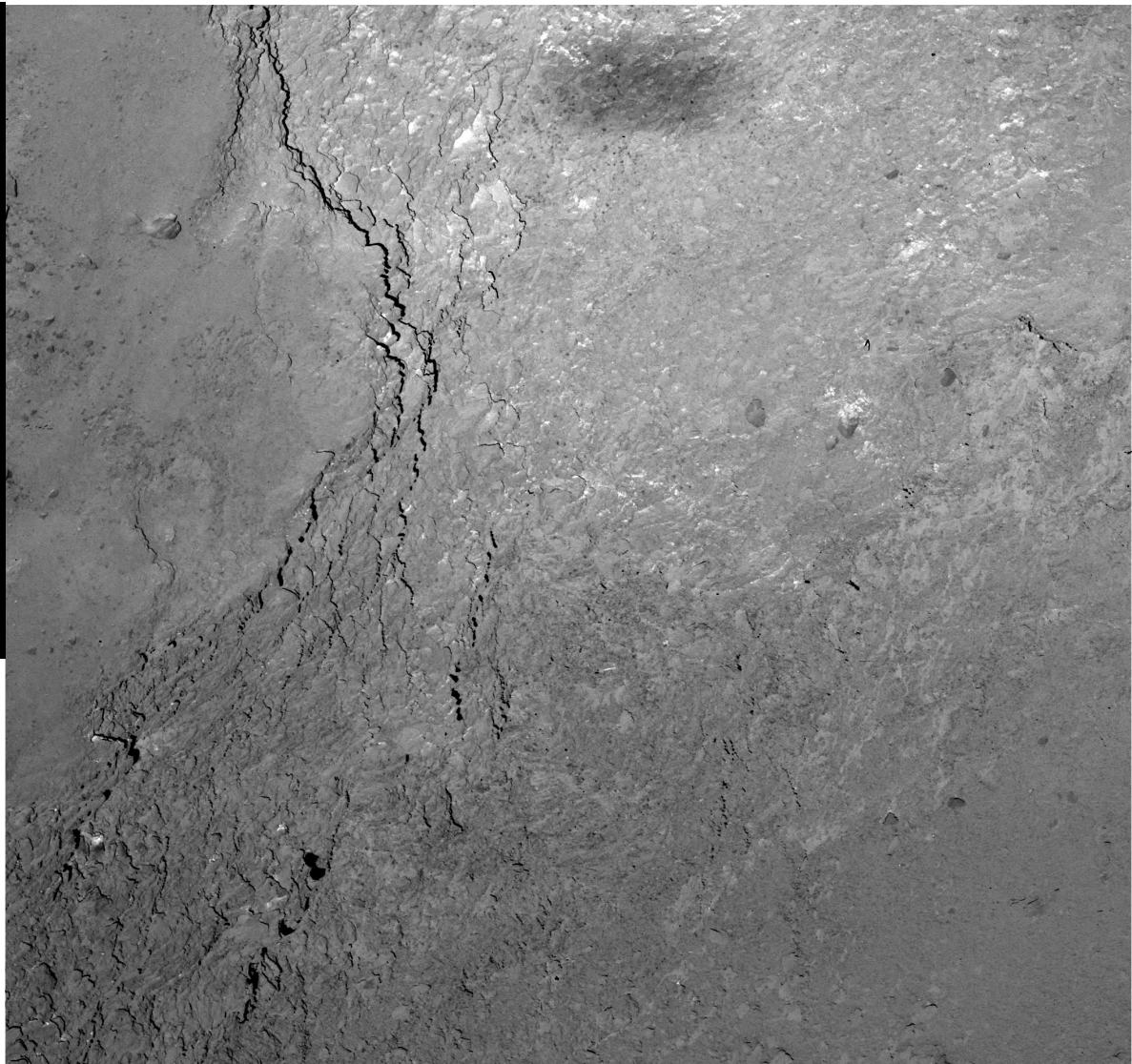
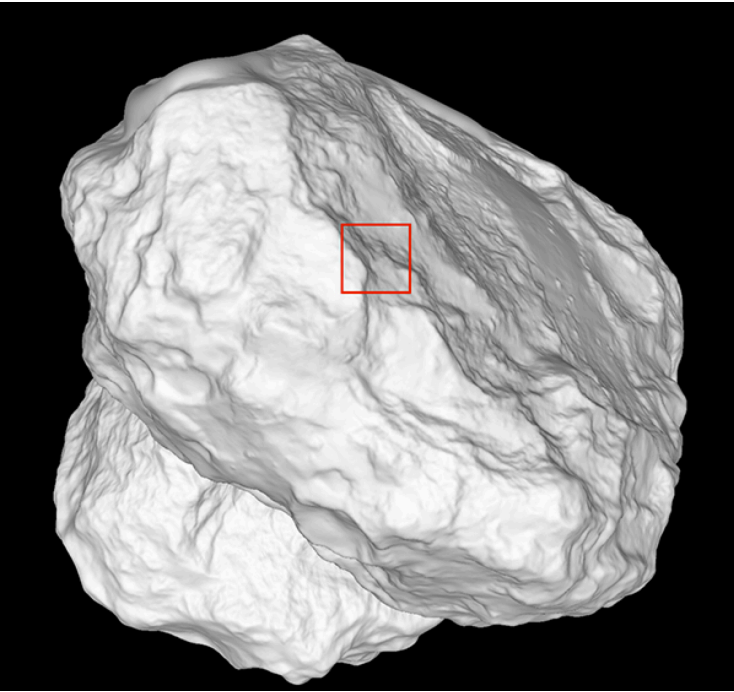


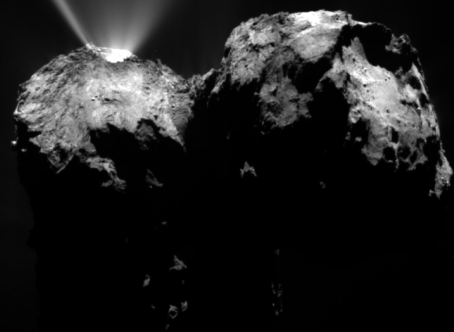
Close flyby



February 14, 2015 at 6 km

(Res.= 11 cm) **Activity everywhere**





Perihelion 13 Agust 2015